Lecture I Hitoshi Murayama (Berkeley) Cargèse, 31/07/2007

Title given by Geraldine:

Supersymmetry: motivations, phenomenology, dark matter candidates, neutrinos

many big subjects!

Plan

1: Non-technical Overview of Supersymmetry what SUSY is supposed to give us 2: From formalism to the MSSM Global SUSY formalism, Feynman rules, soft SUSY breaking, MSSM how to break SUSY, mediation mechanisms 3: SUSY in the universe dark matter, dark energy, problems 4: Neutrino physics oscillation data, implications

Standard Model

A Long History

Since Fermi and Yukawa to the "Standard Model," it took almost 40 years to build

Since deep inelastic scattering and J/ψ to precision measurements, it took almost 30 years to test

 Now many beautiful experimental tests: LEP, Tevatron, Møller scattering, neutrino DIS, atomic parity violation, etc etc

Yet not completely established

Renormalizable Quantum Field Theory

O SU(3)_cxSU(2)_LxU(1)_Y gauge theory

	Q	d	u	L	e	B	W	g	Η	G
SU(3) _C	3	3	3	1	1	1	1	8	1	1
SU(2) _L	2	1	1	2	1	1	3	1	2	1
U(1) _Y	+1/6	-1/3	+2/3	-1/2	+1	0	0	0	-1/2	0
spin	-1/2	+1/2	+1/2	-1/2	+1/2	1	1	1	0	2
flavor	3	3	3	3	3	1	1	1	1	1
seen?	Y	Y	Y	Y	Y	Y	Y	Y	Ν	Ν

Anomaly Cancellation $3 \times 2\left(\frac{1}{6}\right)^3 + 3 \times \left(-\frac{2}{3}\right)^3 + 3 \times \left(\frac{1}{3}\right)^3 + 2\left(-\frac{1}{2}\right)^3 + (+1)^3 = 0$ U(1)³ **U(1)(gravity)**² $3 \times 2\left(\frac{1}{6}\right) + 3 \times \left(-\frac{2}{3}\right) + 3 \times \left(\frac{1}{3}\right) + 2\left(-\frac{1}{2}\right) + (+1) = 0$ $3 \times 2\left(\frac{1}{6}\right) + 2\left(-\frac{1}{2}\right) = 0$ $U(1)(SU(2))^{2}$ $3 \times 2\left(\frac{1}{6}\right) + 3 \times \left(-\frac{2}{3}\right) + 3 \times \left(\frac{1}{3}\right) = 0$ $U(1)(SU(3))^2$ (SU(3))³ $#3 - #3^* = 2 - 1 - 1 = 0$ $(SU(2))^3$, $(SU(3))^2SU(2)$, $SU(3)(SU(2))^2$ $\left(\right)$ SU(2) #2 = 3 + 1 = 4 = evenNon-trivial connection between q & l

General

The most general renormalizable Lagrangian with the given particle content $\mathcal{L} = -\frac{1}{4g'^2} B_{\mu\nu} B^{\mu\nu} - \frac{1}{4g^2} W^a_{\mu\nu} W^{\mu\nu a} - \frac{1}{4g_s^2} G^a_{\mu\nu} G^{\mu\nu a}$ $+\bar{Q}_i i \not D Q_i + \bar{u}_i i \not D u_i + \bar{d}_i i \not D d_i + \bar{L}_i i \not D L_i + \bar{e}_i i \not D e_i$ $+|D_{\mu}H|^{2}+Y_{\mu}^{ij}\bar{Q}_{i}u_{j}\tilde{H}+Y_{d}^{ij}\bar{Q}_{i}d_{j}H+Y_{l}^{ij}\bar{L}_{i}e_{j}H$ $-\lambda (H^{\dagger}H)^{2} + \lambda v^{2}H^{\dagger}H + \frac{\theta}{64\pi^{2}}\varepsilon^{\mu\nu\rho\sigma}G^{a}_{\mu\nu}G^{a}_{\rho\sigma}$

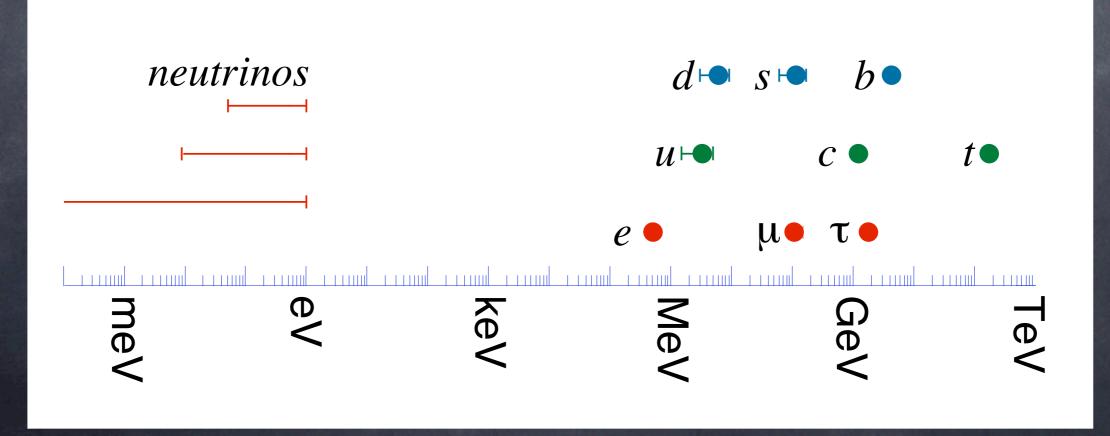
Parameters

 \oslash 3 gauge coupling constants + θ_{QCD} \odot 2 parameters in the Higgs potential (G_F, m_H) $\mathcal{L} = -\frac{1}{4\varrho'^2} B_{\mu\nu} B^{\mu\nu} - \frac{1}{4\varrho^2} W^a_{\mu\nu} W^{\mu\nu a} - \frac{1}{4\varrho^2} G^a_{\mu\nu} G^{\mu\nu a}$ $+\bar{Q}_{i}i\mathcal{D}Q_{i}+\bar{u}_{i}i\mathcal{D}u_{i}+\bar{d}_{i}i\mathcal{D}d_{i}+\bar{L}_{i}i\mathcal{D}L_{i}+\bar{e}_{i}i\mathcal{D}e_{i}$ $+|D_{\mu}H|^{2}+Y_{\mu}^{ij}\overline{Q}_{i}u_{j}\widetilde{H}+Y_{d}^{ij}\overline{Q}_{i}d_{j}H+Y_{l}^{ij}\overline{L}_{i}e_{j}H$ $-\lambda (H^{\dagger}H)^{2} + \lambda v^{2}H^{\dagger}H + \frac{\theta}{64\pi^{2}} \varepsilon^{\mu\nu\rho\sigma}G^{a}_{\mu\nu}G^{a}_{\rho\sigma}$ g'~0.36, g~0.65, g_s~1.2 $G_{F} \sim (300 \text{ GeV})^{-2}$, m_H unknown, $\theta_{QCD} < 10^{-10}$

Parameters

3x3 complex Y^{ij}, Y^{jj}, Y^{ij}, Y^{ij}: 54 real params \odot reparameterization U(3)_{Q,d,u,L,e}⁵/U(1)_{B,e,µ,t}⁴=41 $= -\frac{1}{4\varrho'^2} B_{\mu\nu} B^{\mu\nu} - \frac{1}{4\varrho^2} W^a_{\mu\nu} W^{\mu\nu a} - \frac{1}{4\varrho^2} G^a_{\mu\nu} G^{\mu\nu a}$ $+\bar{Q}_i i \not D Q_i + \bar{u}_i i \not D u_i + \bar{d}_i i \not D d_i + \bar{L}_i i \not D L_i + \bar{e}_i i \not D e_i$ $+|D_{\mu}H|^{2}+Y_{\mu}^{ij}\overline{Q}_{i}u_{j}\widetilde{H}+Y_{d}^{ij}\overline{Q}_{i}d_{j}H+Y_{l}^{ij}\overline{L}_{i}e_{j}H$ $-\lambda (H^{\dagger}H)^{2} + \lambda v^{2}H^{\dagger}H + \frac{\theta}{64\pi^{2}} \varepsilon^{\mu\nu\rho\sigma}G^{a}_{\mu\nu}G^{a}_{\rho\sigma}$ $54-41=13=3_u+3_d+3_l+(3+1)_{CKM}$

Masses and Mixings



Standard Model is extreeemely successful

Take Particle Data Group "Reviews of Particle Physics" with 400+ pages

- With only a few exceptions, all numbers in the book are consistent with the Standard Model with suitably chosen 19 parameters
- Some of them tested at $10^{-9} 10^{-12}$ level

Standard Model is extreeemely successful

- flavor approximately conserved (apart from small mixing in V_{СКМ})

So, what's the problem?

empirically incomplete

neutrino mass

ø dark matter

ø dark energy

nearly scale-invariant apparently acausal density fluctuation

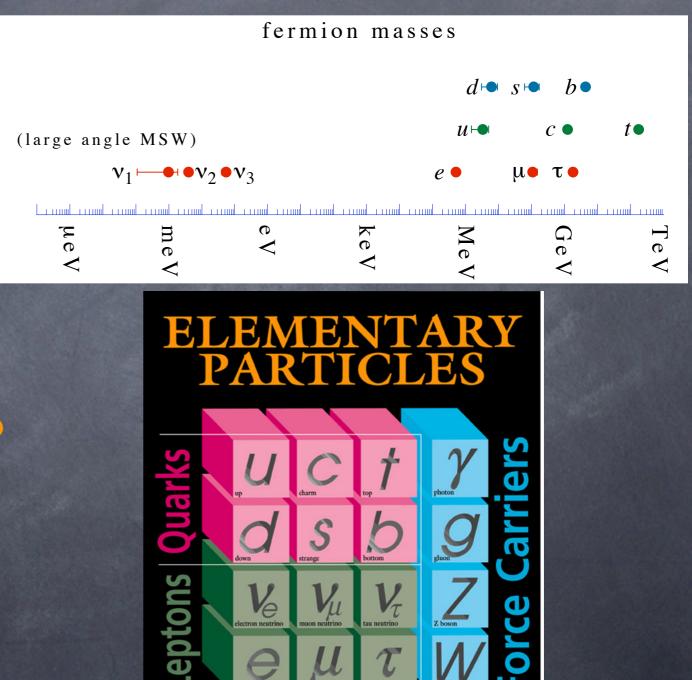
ø baryon asymmetry

aesthetically unacceptable

structure is quite complicated
many naturalness problems
no quantum gravity
questions in four categories

Big Questions -Horizontal-

- Why are there three generations? What physics determines the pattern of masses and mixings? Why do neutrinos have mass yet so light? What is the origin of CP violation? \odot Why $\theta_{QCD} \ll 10^{-10}$?
- What is the origin of matter anti-matter asymmetry in Universe?



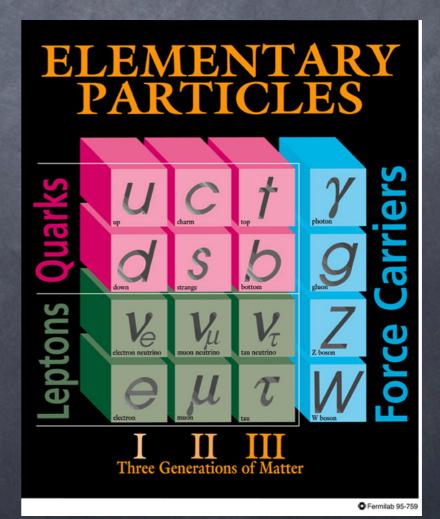
Fermilab 95-759

Big Questions -Vertical-

Why are there three unrelated gauge forces? Why is strong interaction strong? Charge quantization anomaly cancellation quantum numbers Is there a unified description of all forces? Ø Why is m_W≪M_{Pl}?
 (Hierarchy Problem)

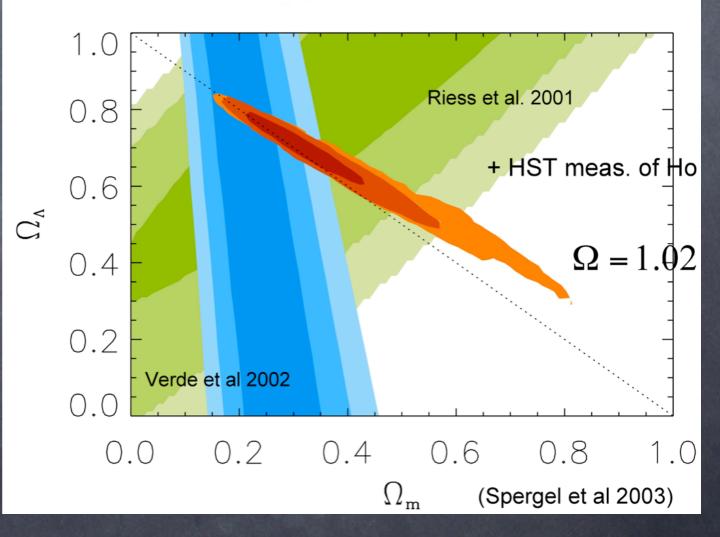
$$Q(\mathbf{3}, \mathbf{2}, +\frac{1}{6}), \quad u(\mathbf{3}, \mathbf{1}, +\frac{2}{3}), \quad d(\mathbf{3}, \mathbf{1}, -\frac{1}{3}),$$

 $L(\mathbf{1}, \mathbf{2}, -\frac{1}{2}), \quad e(\mathbf{1}, \mathbf{1}, -1)$



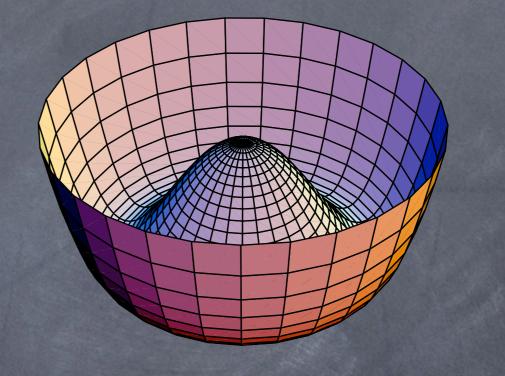
Big Questions -From the Heaven-

- What is Dark Matter? What is Dark Energy? Why now? (Cosmic coincidence problem) What was Big Bang? Why is Universe so big? (flatness problem, horizon problem)
- How were galaxies and stars created?



Big Questions -From the Hell-

- What is the Higgs boson?
- Why does it have negative masssquared?
- Why is there only one scalar particle in the Standard Model?
- Is it elementary or composite?
- Is it really condensed in our Universe?





electric force

weak force

Standard Model is fragile

The minute you allow for additional fields and/or gauge groups, much of the success is destroyed

suppressed flavor-changing neutral currents
no proton decay

o no neutrino mass either (good&bad)

consistency with precise electroweak data

In a stressive CP violation (e/n EDM)

no charge/color breaking

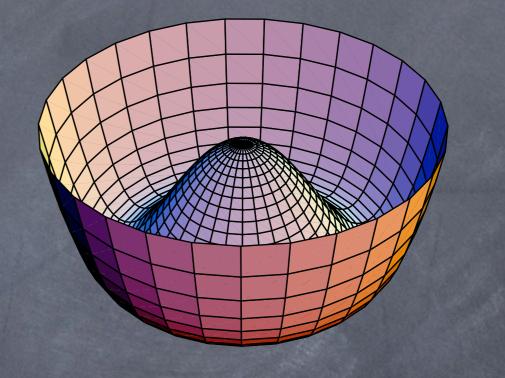
Standard Model is fragile

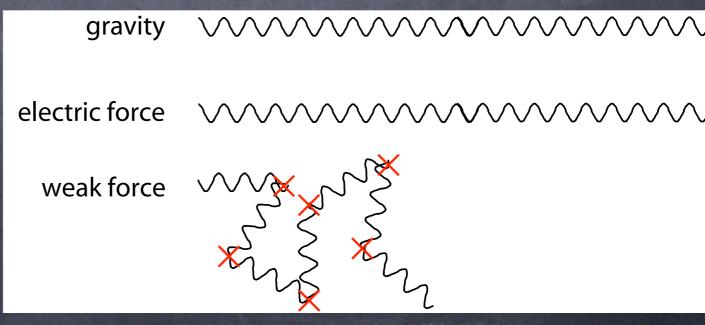
The minute you allow for parameters to vary, it exhibits very different physics \odot take $m_d < m_u$, all protons decay to neutrons and there are no atoms \odot if v>>TeV, $|m_n-m_p|$ too big and neutron decays even inside nuclei; no chemistry, no life If $m_c \sim m_t$, no J/ψ before the end of cold war and no high-energy physics funding by now

Dark Field = cosmic superconductor

Big Questions -From the Hell-

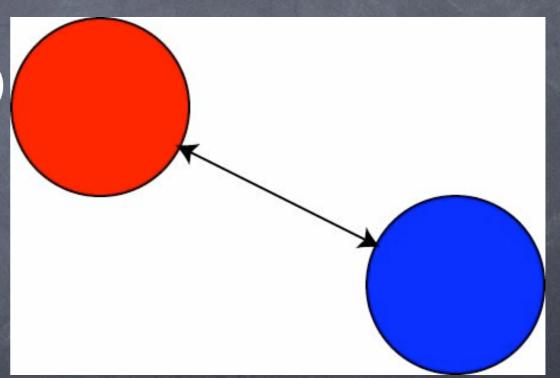
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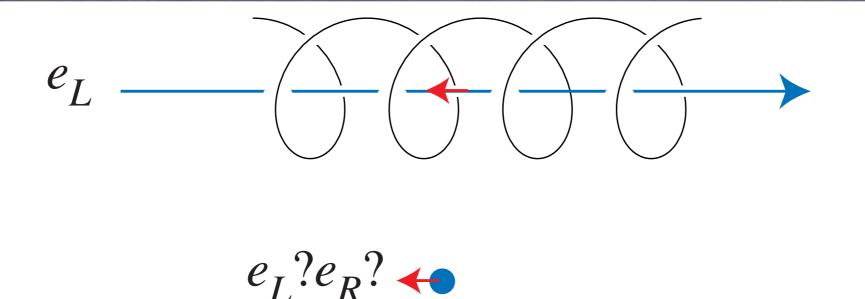


Mystery of the weak force

Gravity pulls two massive bodies (long-ranged) Electric force repels two
 like charges (long-ranged) Weak force pulls protons and electrons (shortranged) acts only over 0.00000001 nanometer [need it for the Sun to burn!] We know the energy scale: 0.3 TeV



Mystery deepens

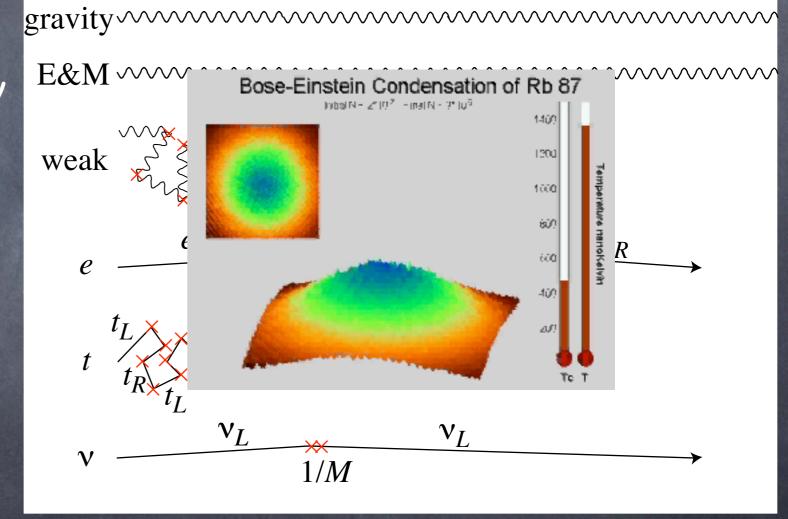


- Suclear beta decay is due to a yet another force, the weak force
- Strangely, only left-handed particles participate in the weak force
- That sounds OK as long as they are moving
- ø but when they stop???

We are swimming in Dark Field

There is quantum liquid filling our Universe

- It doesn't disturb gravity or electric force
- It does disturb weak force and make it shortranged
- It slows down all elementary particles from speed of light
- What is it?? Extremely bizarre theory!



Cosmic Superconductor

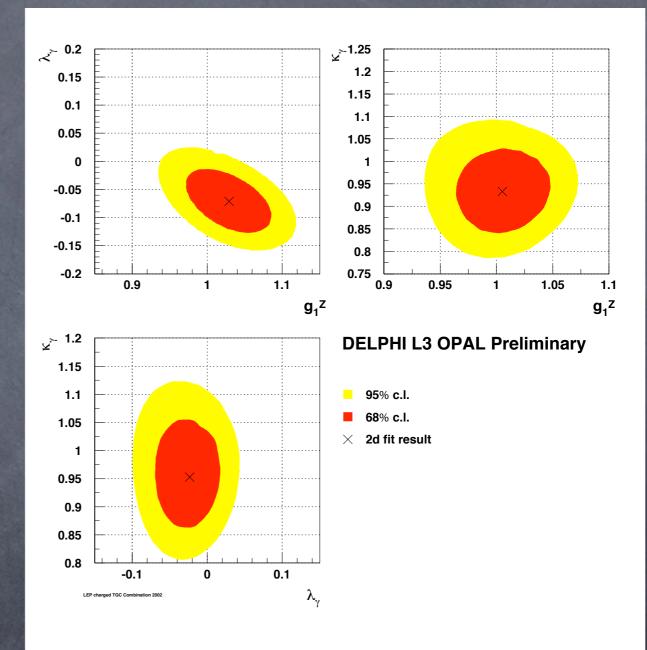
 In a superconductor, magnetic field gets repelled (Meißner effect), and penetrates only over the "penetration length"
 ⇒ Magnetic field is short-ranged!

Imagine a physicist living in a superconductor

- She finally figured:
 - magnetic field must be long-ranged
 - there must be a mysterious charge-two "Dark Field" in her "Universe"
 - But doesn't know what the Dark Field is, nor why it is there
 - Doesn't have enough energy (gap) to break up Cooper pairs

Textbook

W and Z are massive vector bosons Only known consistent (renormalizable) quantum field theory of massive vectors is gauge theory with Higgs mechanism Therefore, W and Z bosons must be gauge bosons, broken by a Higgs

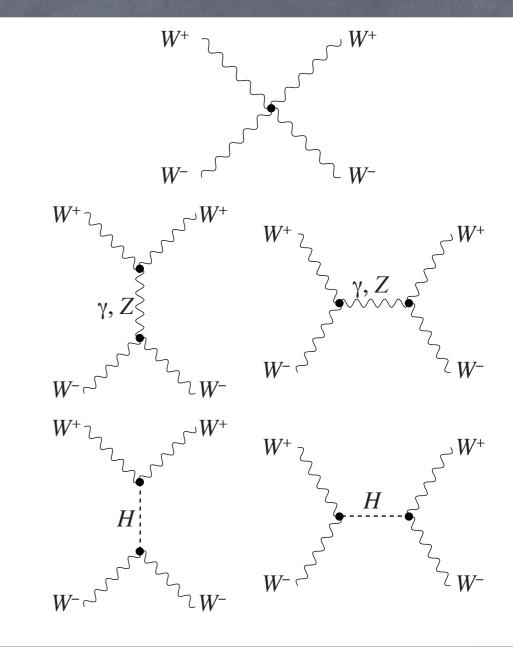


Now added evidence of vector boson self-couplings

unitarity

 W-boson scattering grows with energy
 A~G_FE² and violates

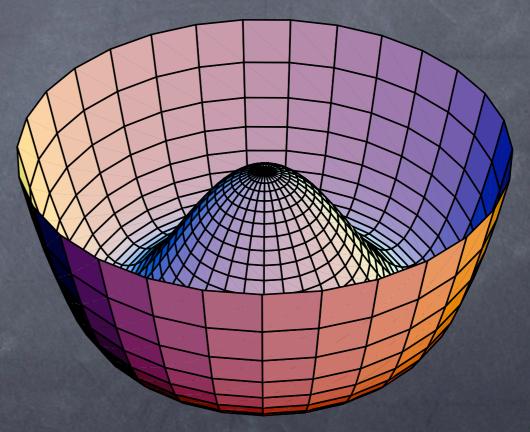
unitarity at 1.8TeV
If you allow only one extra particle beyond what we know to restore unitarity, the only possibility is to add a spin zero particle whose couplings are precisely those of the SM Higgs



C. H. Llewellyn Smith; D. A. Dicus and V. S. Mathur; J. M. Cornwall, D. N. Levin and G. Tiktopoulos

ugly

- $V = \lambda |H|^4 \mu^2 |H|^2$
- Why negative masssqured?
- Why only one scalar in the SM?
- Hierarchy problem
 because of its quadratic
 divergence
- does not appear fundamental, i.e.
 Ginzburg-Landau vs BCS

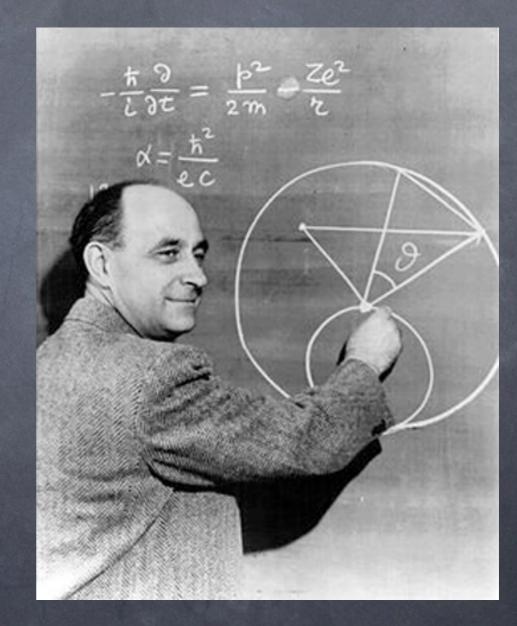


Fermi's dream era

Fermi formulated the first theory of the weak force (1933)

The required energy scale to study the problem known since then: ~TeV

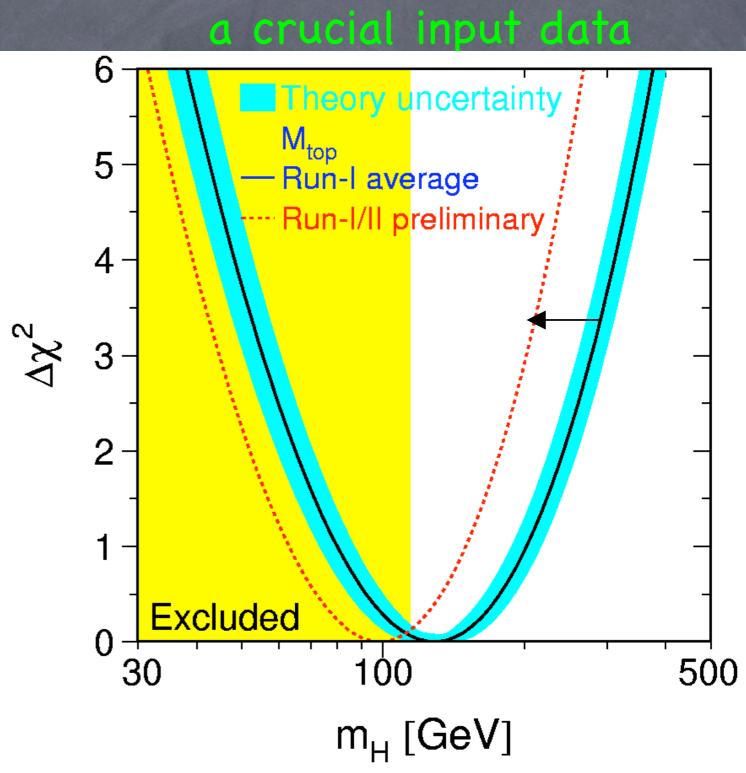
We are finally getting there!



Gap Excitation

the top mass is

We know the energy scale of the problem: **G**_{**r**}≈(300 GeV)⁻² the gap excitation is called "Higgs boson" Ourrent data
 Ourrent
 Ourrent data
 Ourrent
 combined with the Standard Model theory predict m_µ<208GeV (95%CL)



Kick out Dark Field from the vacuum

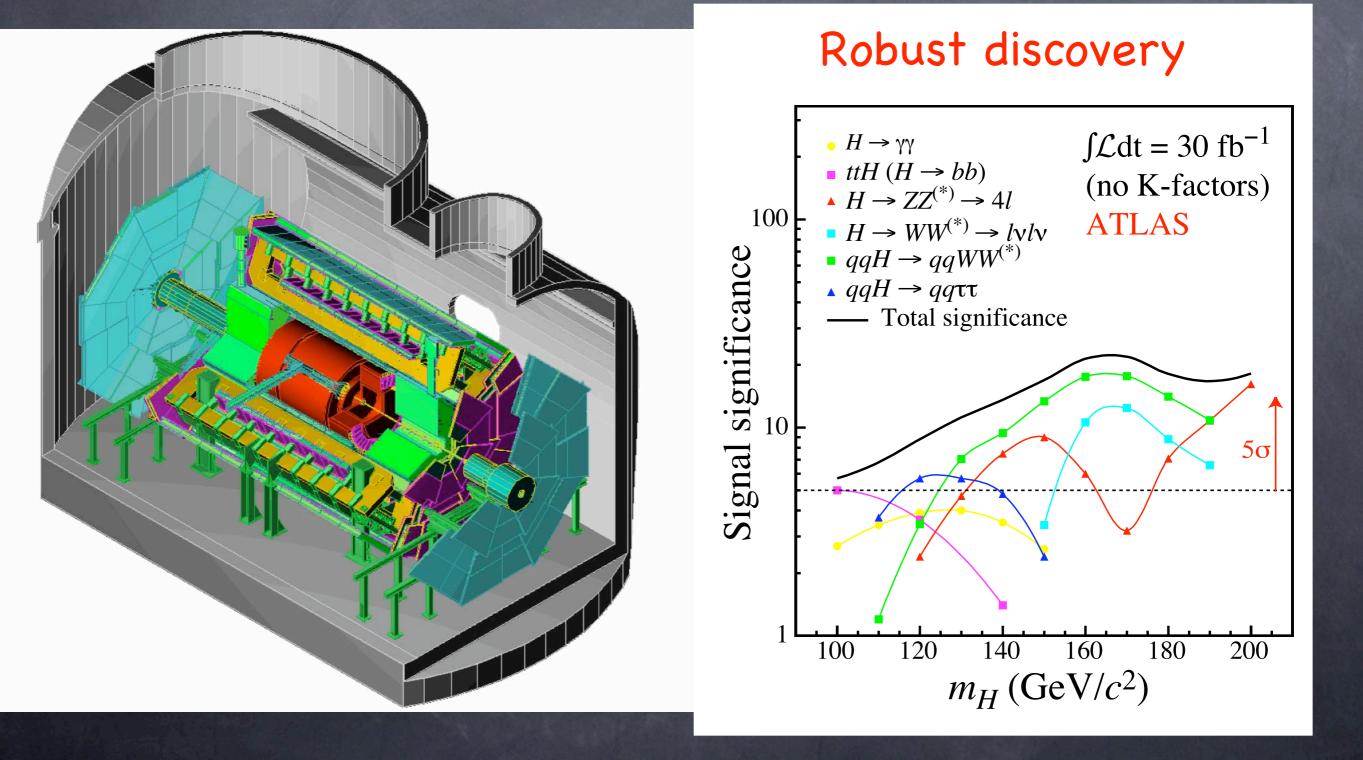
We know the energy scale of the problem:

0.3 TeV

- pump energy into empty space to kick out whatever makes Dark Field: "Higgs boson"
- LHC will find it!!!!!



Higgs at ATLAS



Better be sure

- For something this bizarre, we'd better make sure
- Is the particle discovered really the Higgs boson?
 - Is it really responsible for particle masses?
 - Does this have the right properties?
 Is it really stuck in our Universe?
 Need detailed measurements for the proof

Questions to be answered

Is the particle discovered really the Higgs boson?

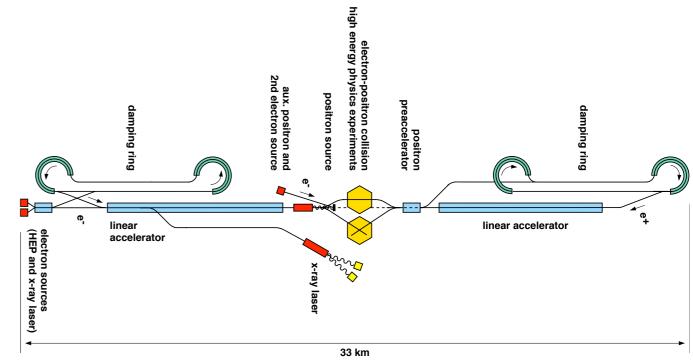
- Is it really responsible for particle masses?
- Does this have the right quantum number J^p=0+?

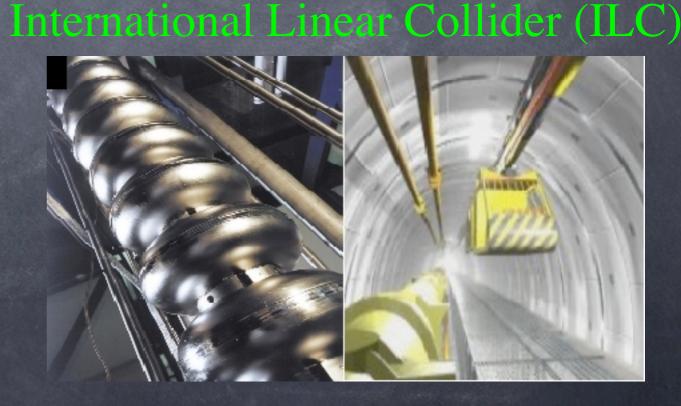
Is it condensed in the Universe, namely here?
Prove it is the "Mother of Mass"

- \oslash Coupling \propto mass?
- Size of the condensate

Linear Collider

- Selectron-positron collider
- e⁻, e⁺ point-like with no structure
- Well-understood environment
- Linear instead of ring to avoid synchrotron loss
- Super-high-tech machine
- Accelerate the beam over >15km
- Focus beam down to a few nanometers and make them collide



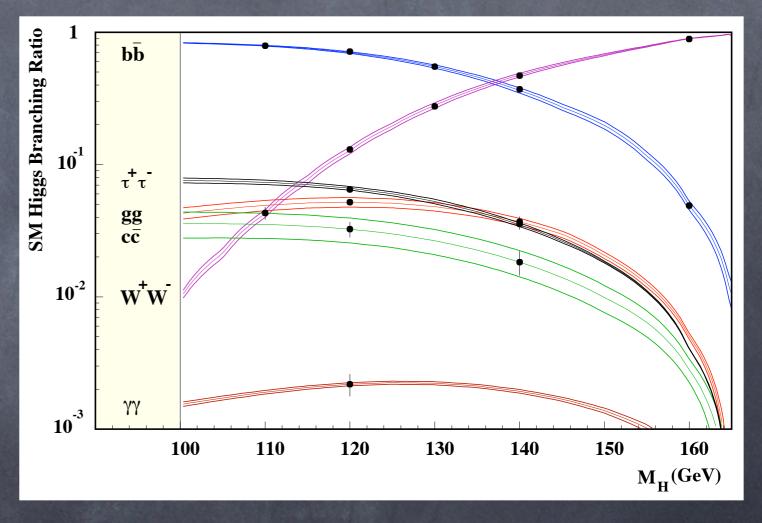


Prove its coupling \propto

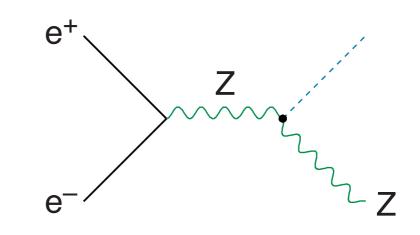
 Branching Fractions test the relation

coupling ∝ mass

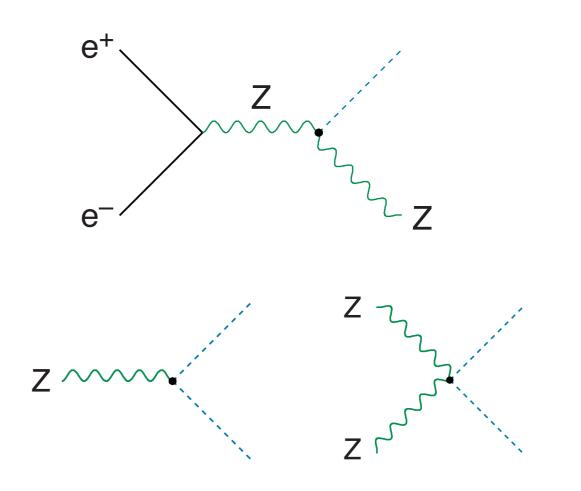
⇒ proves that Higgs
 Boson is the Mother of Mass



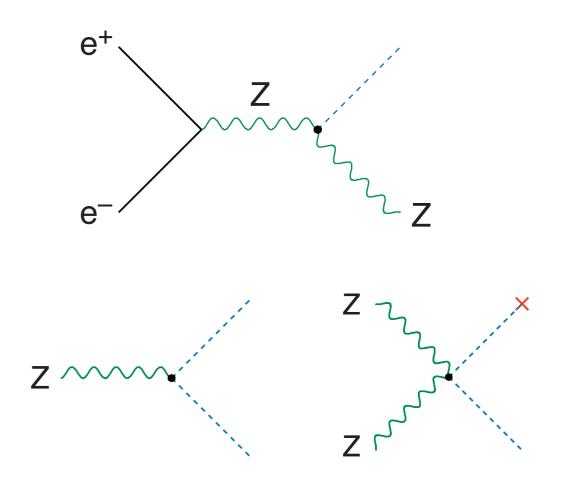
ZH final stateProve the ZZH vertex



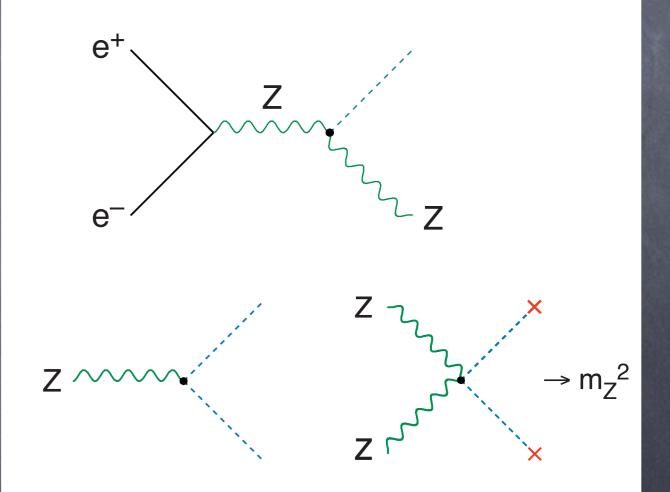
⊘ ZH final state
⊘ Prove the ZZH vertex
⊘ We know Z:gauge boson, H: scalar boson
⇒ only two types of vertices



Prove the ZZH vertex We know Z:gauge boson, H: scalar boson \Rightarrow only two types of vertices Need a condensate to get ZZH vertex \Rightarrow proves it is condensed in Universe



Prove the ZZH vertex We know Z:gauge boson, H: scalar boson \Rightarrow only two types of vertices Need a condensate to get ZZH vertex \Rightarrow proves it is condensed in Universe \Rightarrow proves it is the origin of m_Z HM, hep-ex/9606001

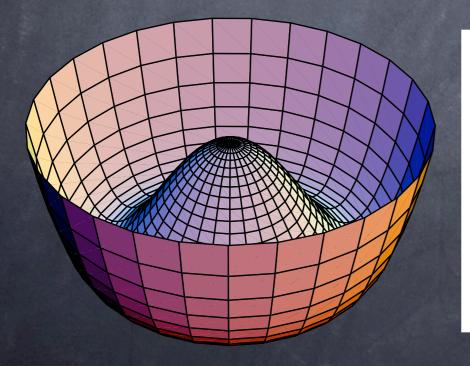


Post-Higgs Problem

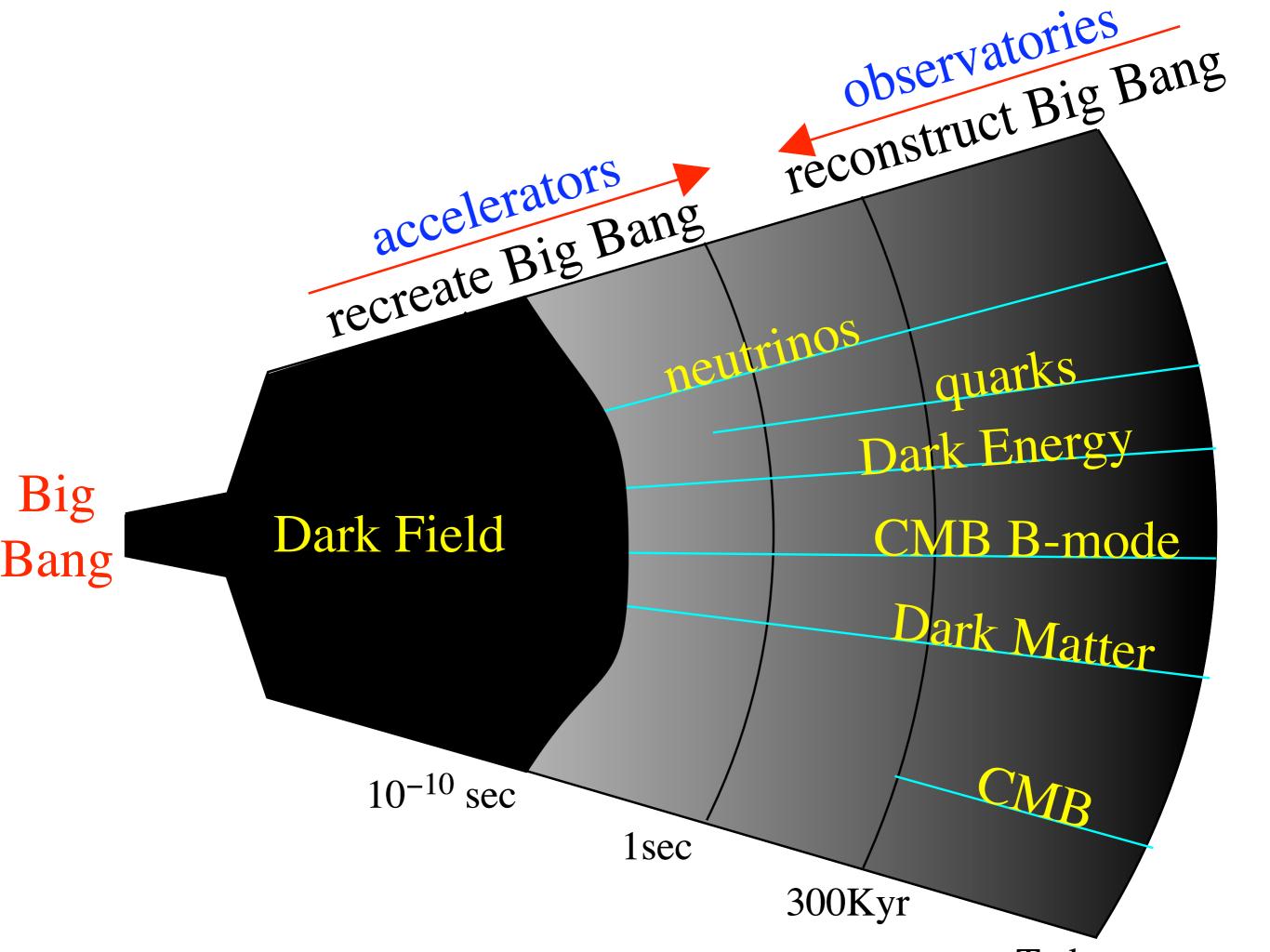
We see "what" is stuck in our universe
But we still don't know "why"
Two problems:
Why anything is condensed at all
Why is the scale of Dark Field 0.3TeV much much smaller than the scale of gravity ~10¹⁵ TeV

Second Explanation most likely to be at ≤TeV scale because this is the relevant energy scale

Hierarchy Problem

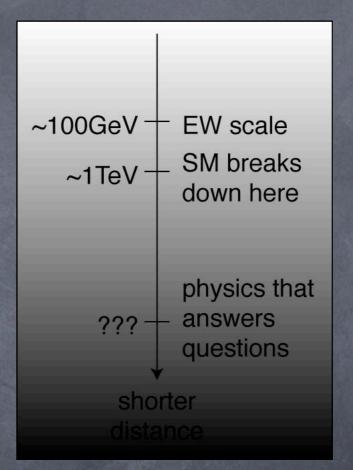


gravity weak force



The Main Obstacle

- We look for physics beyond the Standard Model that answers big questions about early universe
- By definition, that is physics at shorter distances
- Then the Standard Model must survive down to whatever shorter distance scale
- ✓ Hierarchy problem is the main obstacle to do so
 ⇒ We can't even get started!



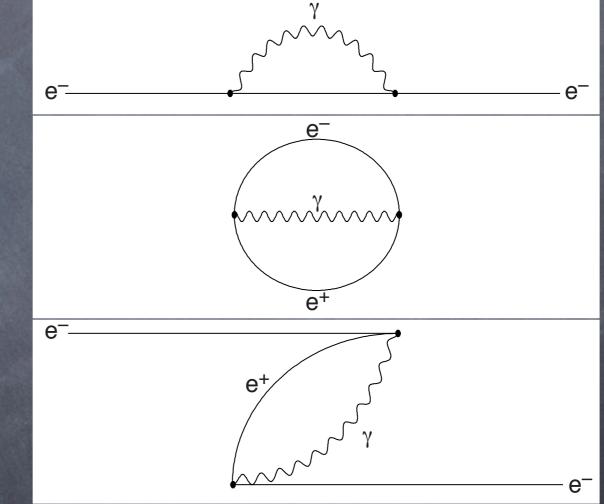
Once upon a time, there was a hierarchy problem...

- At the end of 19th century: a "crisis" about electron
 - Solution Like charges repel: hard to keep electric charge in a small pack
 - Selectron is point-like
- Need a lot of energy to keep it small!

 $\Delta m_e c^2 \sim \frac{e^2}{r_e} \sim \text{GeV} \frac{10^{-17} \text{cm}}{r_e}$ Correction $\Delta m_e c^2 > m_e c^2$ for $r_e < 10^{-13} \text{cm}$ Breakdown of theory of electromagnetism \Rightarrow Can't discuss physics below 10^{-13}cm

Anti-Matter Comes to Rescue by Doubling of #Particles

 Electron creates a force to repel itself Solution Vacuum bubble of matter anti-matter creation/annihilation Selectron annihilates the positron in the bubble \Rightarrow only 10% of mass even for Planck-size $r_e \sim 10^{-33}$ cm

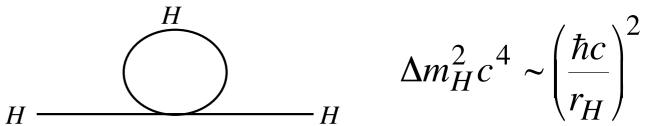


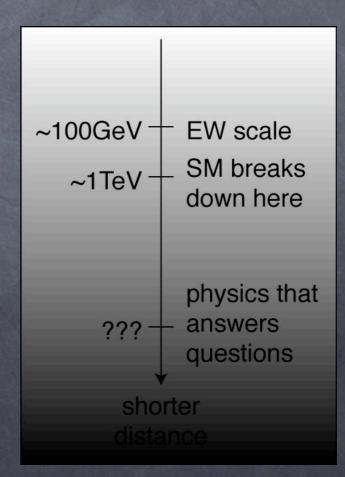
 $\Delta m_e \sim m_e \frac{\alpha}{4\pi} \log(m_e r_e)$

Higgs repels itself, too

 Just like electron repelling itself
 because of its charge, Higgs boson also repels itself
 Requires a lot of

- energy to contain itself in its point-like size!
- Breakdown of theory of weak force
- © Can't get started!

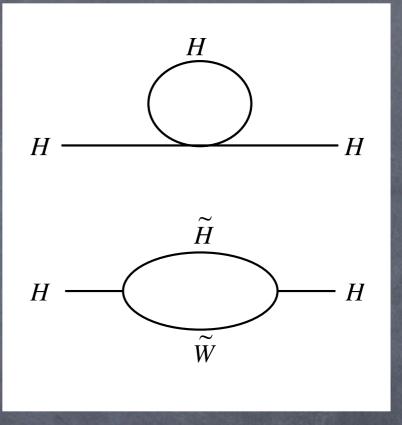




History repeats itself?

⊘ Double #particles
 again ⇒

superpartners Vacuum bubbles" of superpartners cancel the energy required to contain Higgs boson in itself Standard Model made consistent with whatever physics at shorter distances

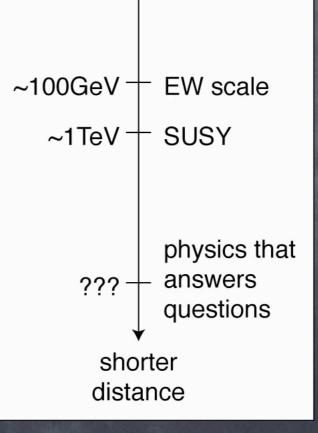


 $\Delta m_H^2 \sim \frac{\alpha}{4\pi} m_{SUSY}^2 \log(m_H r_H)$

Opening the door

Once the hierarchy problem solved, we can get started to discuss physics at shorter distances and earlier universe.
 It opens the door to the next level:
 Hope to answer big questions

The solution to the hierarchy problem itself, e.g., SUSY, provides additional probe to physics at short distances



Three Directions

History repeats itself

- Crisis with electron solved by anti-matter

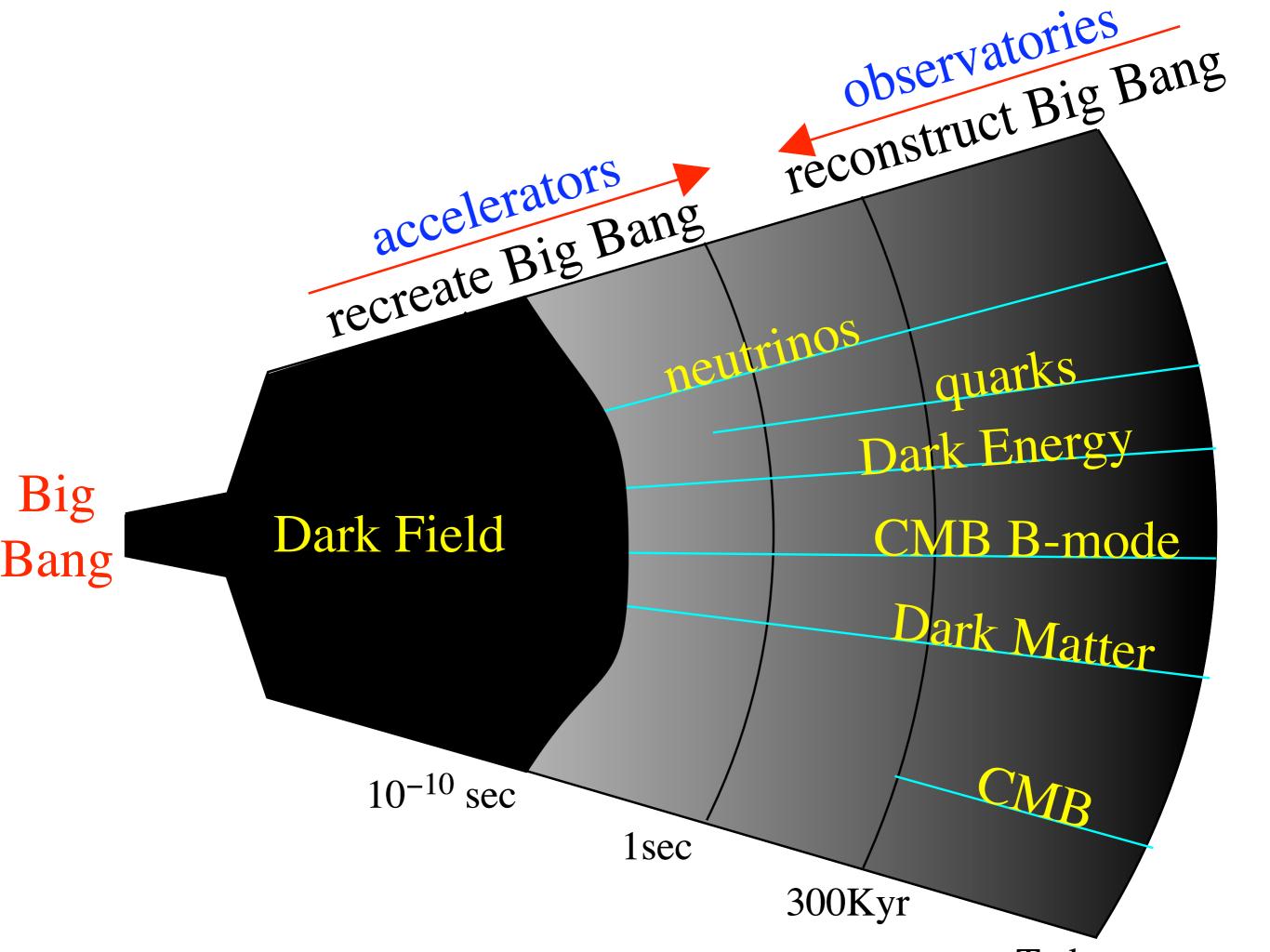
Learn from Cooper pairs

- Cooper pairs composite made of two electrons
- Higgs boson may be fermion-pair composite
 technicolor

Physics as we know it ends at TeV Iltimate scale of physics: quantum gravity May have quantum gravity at TeV ⇒ hidden dimensions (0.1 mm to 10⁻¹⁷ cm)

More Directions

Higgs boson as a Pseudo-Nambu-Goldstone boson (Little Higgs)
Higgs boson as an extra-dimensional gauge boson (Gauge-Higgs Unification)
Fat Higgs (Composite)
Higgsless and W[±] as Kaluza-Klein boson
technicolorful supersymmetry



Anthropic reason?

- Maybe electroweak scale is what it is simply because that is the only kind of universe we can observe
- if v>>TeV, |m_n-m_p| too big and neutron decays even inside nuclei; no chemistry, no life
- after all, cosmological constant is "clearly" not natural
- If this were the case, there may be nothing else to be seen at the LHC other than the "elementary" Higgs boson
- See, however, "A Universe Without Weak Interactions" (Harnik, Kribs, Perez)

Growing Uneasiness

If there are indeed new physics ≤TeV that makes EW scale "natural" They why didn't it show any imprints in EW precision observables? 𝔅 FCNC (K⁰-K⁰, B_d→φK_s, B_s-B_s)? OP violation (EDMs)? Maybe no new physics ≤TeV?

uneasiness in cosmology

B

- Before COBE, upper limit on CMB anisotropy kept getting better and better
- Before 1998, the
 second s universe appeared younger than oldest stars
- cosmologists got antsy
- Crisis in standard cosmoloay"

T (uK) Angular scale Ba 2° 0.5° 90° 0.2° 6000 F Sho 5000 $\begin{matrix} l(l+1)C_l / 2\pi & [\mu K^2] \\ 1000 & 000 \\ 0000 & 000 \\$ A ne rewr univ 1000 Tim 0 500 1000 100 10 Multipole moment *l*

De.

2W

les

EWSB models

We know the energy scale of electroweak symmetry breaking; we are getting there! We need to understand what kind of "quantum liquid" we are dipped in why it is there what causes it condense Otherwise we can't study physics at shorter distances, universe before the BEC

Supersymmetry

Supersymmetry

SUSY Higgs only one of many scalars that happen to acquire negative mass-squared SUSY stabilizes the hierarchy easily consistent with the EW precision observables because it is "decoupling" physics In fully consistent, renormalizable, calculable theory

can be connected to GUT, string, etc

A Broken Symmetry

- Supersymmetry predicts boson and fermion to have the same mass
- Clearly not true in nature
- It has to be broken, partners heavier than the SM particles
- Once broken, it is natural for partners to be heavier as their masses allowed by SU(2)×U(1), while quark, lepton, W, Z masses forbidden

Soft supersymmetry breaking

- Purpose of supersymmetry is to protect hierarchy
- Arbitrary terms in Lagrangian that break supersymmetry reintroduce power divergences
- Soft supersymmetry breaking" classified: mλλ, m²_{ij}φ^{i*}φ^j, A_{ijk}φ^jφ^jφ^k, B_{ij}φ^jφ^j, C_iφ^j
 gaugino mass, squark/slepton mass-squared, etc
 Dark horse terms (not always allowed): φ^{j*}φ^jφ^k, λψ^j, ψⁱψ^j

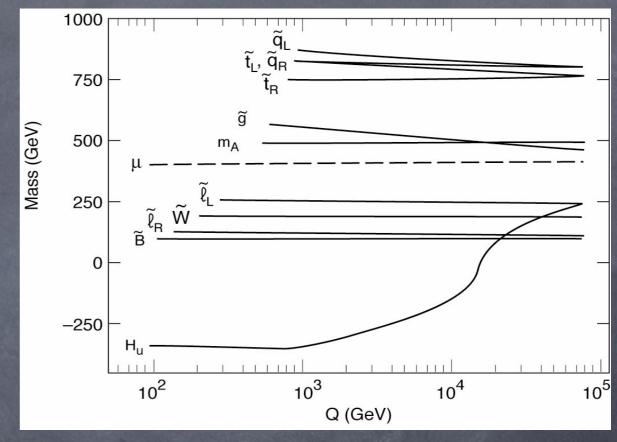
Radiative Symmetry Breaking

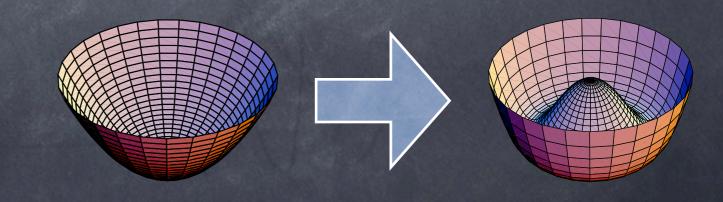
(Inoue et al; Alvarez-Gaumé et al; Ibañez-Ross)

64

 In the MSSM, electroweak symmetry does not get broken
 Only after supersymmetry is broken, Higgs can obtain a VEV v~m_{susy}

EWSB is as a consequence of supersymmetry breaking
 EW symmetry and hierarchy "protected" by supersymmetry

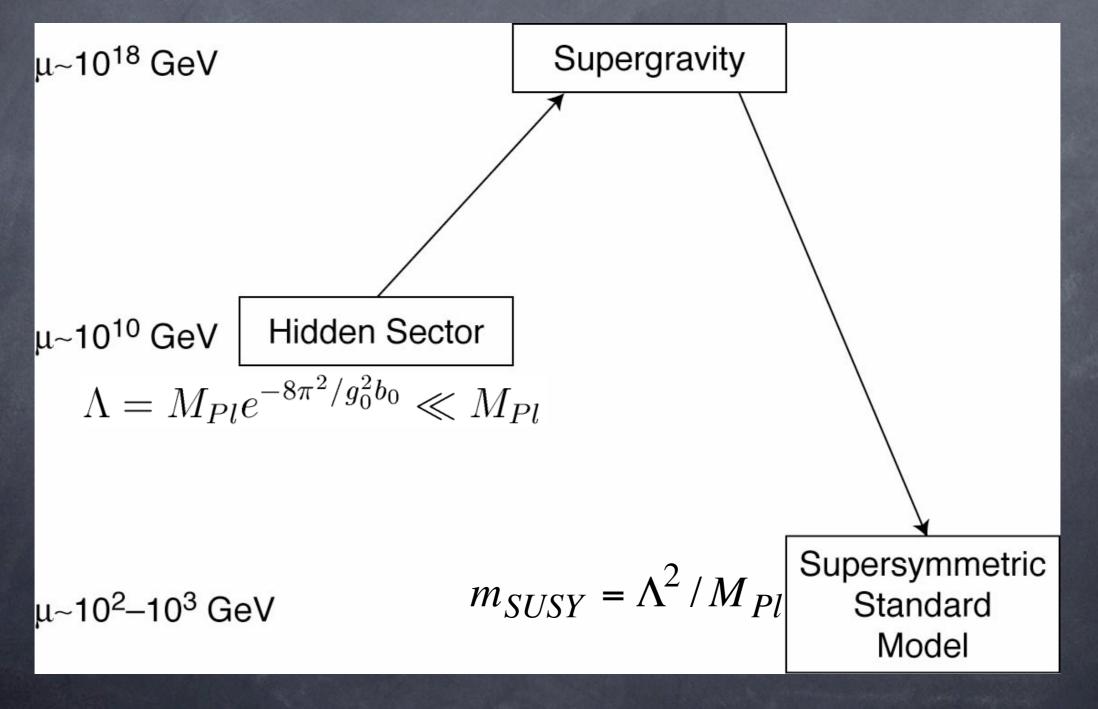




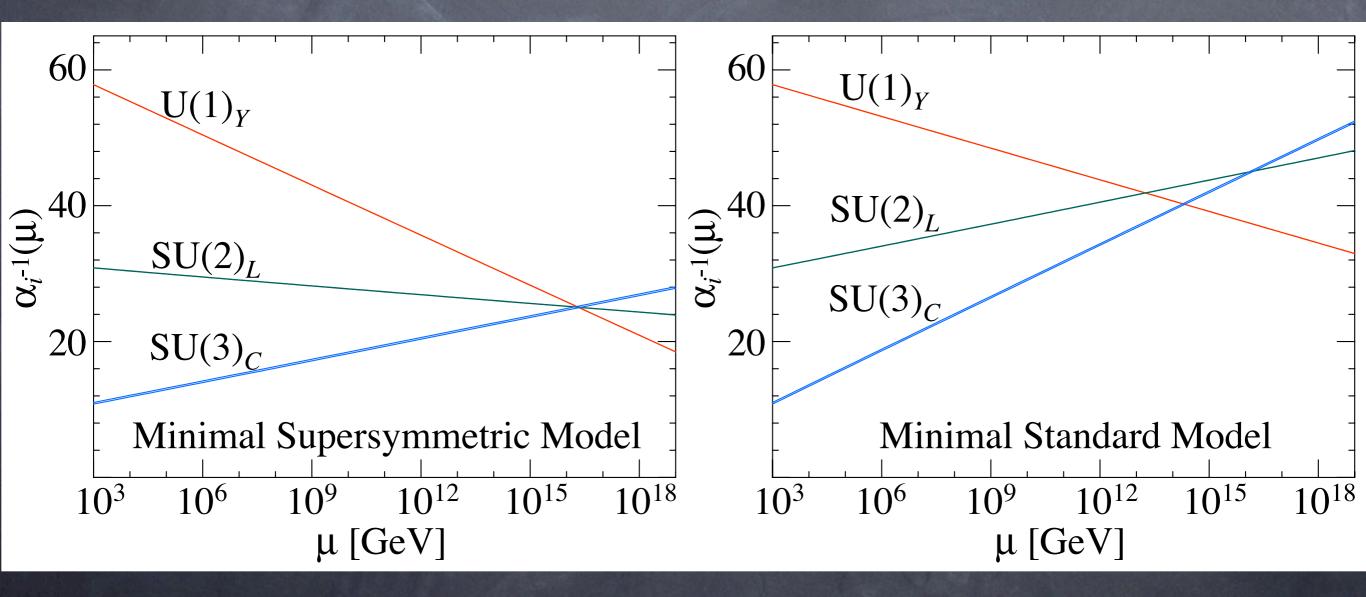
Origin of Hierarchy (Witten)

 $o v \ll M_{\rm Pl}$ because $v \sim m_{\rm SUSY} \ll M_{\rm Pl}$ Ø Why msusy≪Mpl?
 Idea: running coupling constant SUSY broken by strong gauge dynamics with Dynamical supersymmetry breaking" $\Lambda=M_{Pl}e^{-8\pi^2/g_0^2b_0}\ll M_{Pl}$ @ e.g., SO(10) with one 16 (Affleck, Dine, Seiberg; HM)

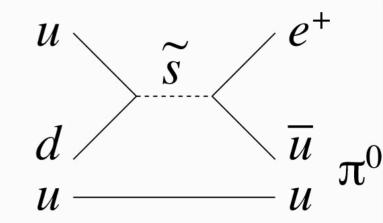
The paradigm



gauge coupling unification



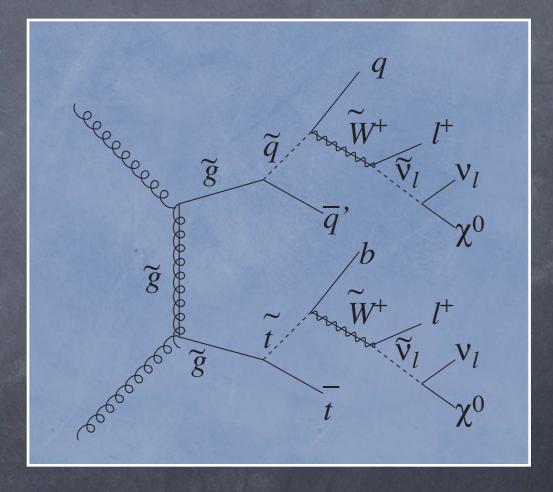




- B, L-conservation not automatic
- W=udd+QdL+LLe+LHu
- \odot If they exist with O(1) couplings:
- $\tau_p \sim m_{sq}^4 / m_p^5 \sim 10^{-12}$ sec!
- Product of two couplings < 10⁻²⁶
- Impose R-parity = $(-1)^{3B+L+2S}$
- Forbids B and L number violation
- \circ R-parity is non-anomalous; may be gauged
- Stable Lightest Supersymmetric Particle
 ⇒ Cold Dark Matter
- SUSY particles always pair-produced and decay into the LSP: missing energy signal

Supersymmetry at hadron colliders

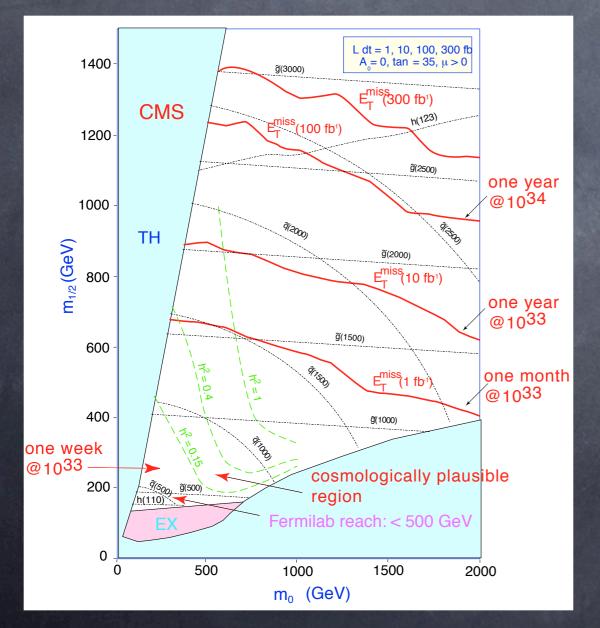
- colored particles
 produced dominantly
- R-parity:
 - ø production in pairs
 - @ decay to LSP
- Solution LSP escapes
 detection: "missing E_T "

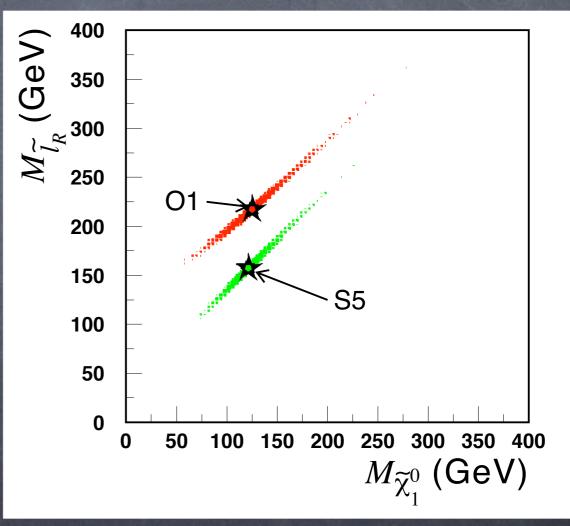


Supersymmetry

Tevatron/LHC will discover

Can do many measurements at LHC



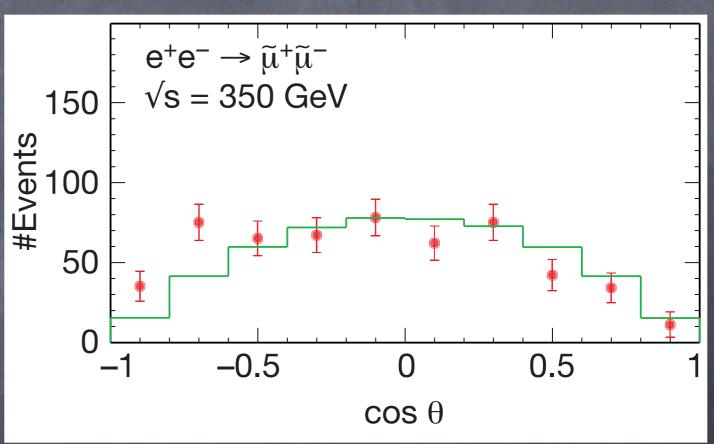


Prove Supersymmetry

 Discovery at Tevatron Run II and/or LHC
 Test they are really superpartners
 Spins differ by 1/2
 Same SU(3)×SU(2) ×U(1) quantum numbers

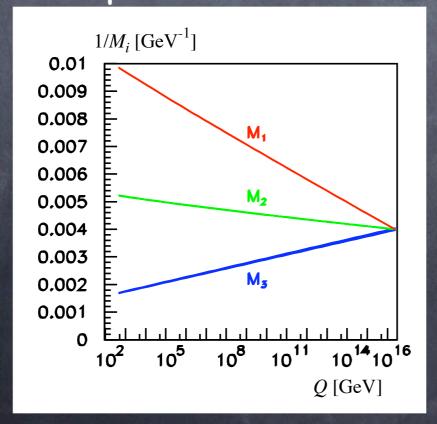
 Supersymmetric couplings

Spin 0?

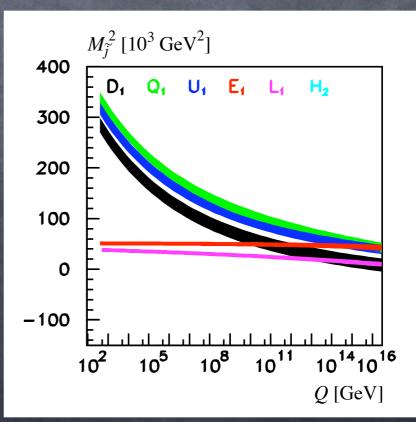


Gaugino and scalars

Gaugino masses test unification indep. of intermediate scales and/or extra multiplets



Scalar masses test beta functions at all scales, depend on the particle content (Kawamura, HM, Yamaguchi)



Particle Dark Matter

 WIMP (Weakly Interacting Massive Particle) strongly favored
 Stable heavy particle produced in early Universe, left-over from near-complete annihilation

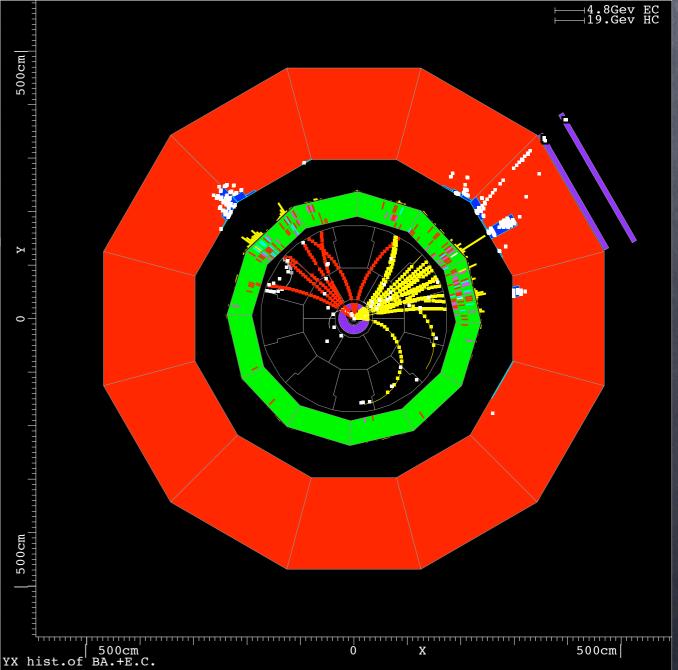
 $\Omega_M = \frac{0.756(n+1)x_f^{n+1}}{g^{1/2}\sigma_{ann}M_{Pl}^3} \frac{3s_0}{8\pi H_0^2} \approx \frac{\alpha^2/(\text{TeV})^2}{\sigma_{ann}}$

TeV the correct energy scale
We hope to produce DM directly at colliders

Producing Dark Matter in the laboratory

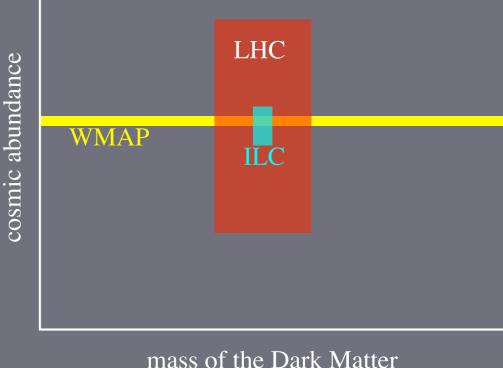
Collision of high-energy particles mimic Big Bang

- We hope to create Dark Matter particles in the laboratory
- Look for events where energy and momenta are unbalanced "missing energy" E_{miss}
- Something is escaping the detector
- ⊘ electrically neutral, weakly interacting
 ⇒ Dark Matter!?

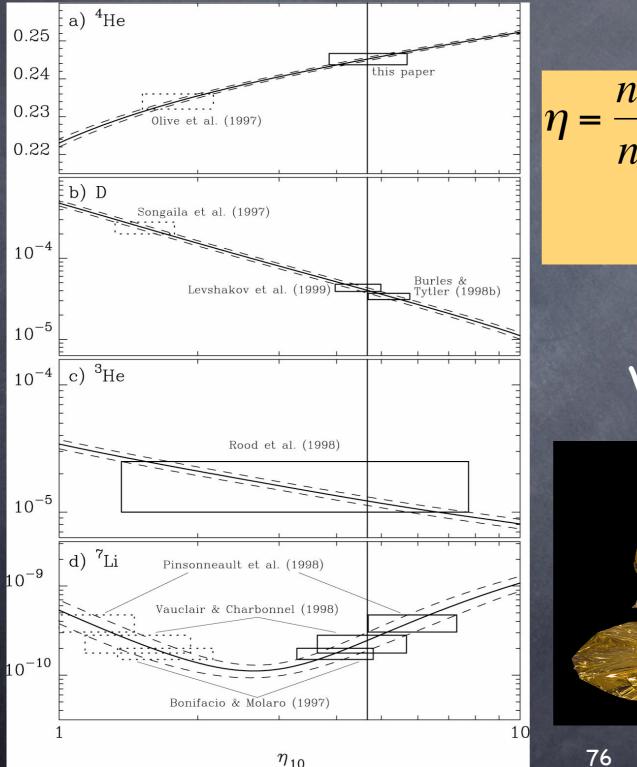


How do we know what Dark Matter is?

cosmological measurement of dark matter \odot abundance $\propto \sigma_{ann}^{-1}$ ø detection experiments cosmic abundance scattering cross section WMAH ø production at colliders mass, couplings can calculate cross sections If they agree with each other: \implies Will know what Dark Matter is \Rightarrow Will understand universe back to t~10⁻¹⁰



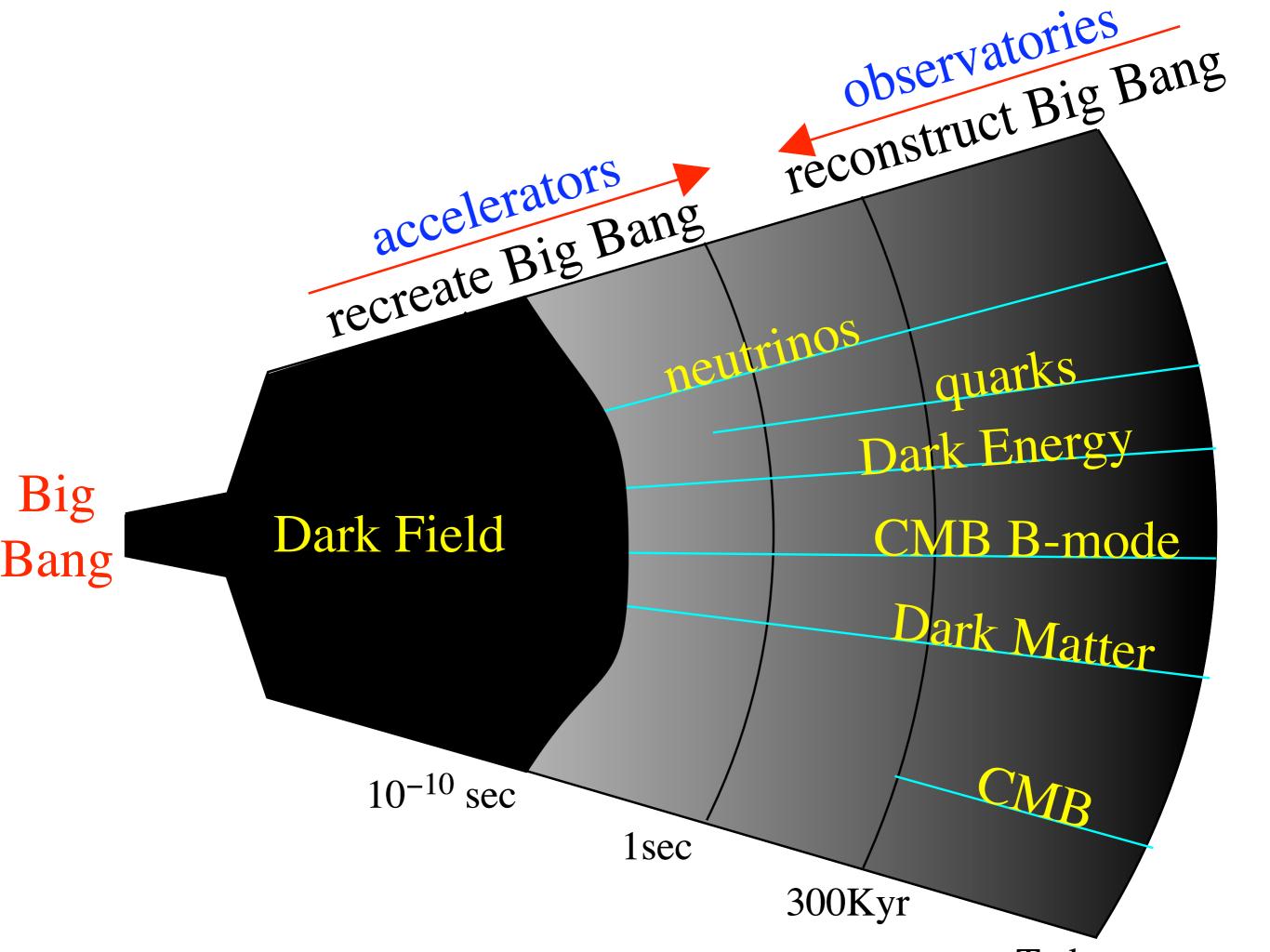
Big-Bang Nucleosynthesis Cosmic Microwave Background



$$\frac{B}{P} = (4.7^{+1.0}_{-0.8}) \times 10^{-10} \text{ (Thuan, Izatov)}$$

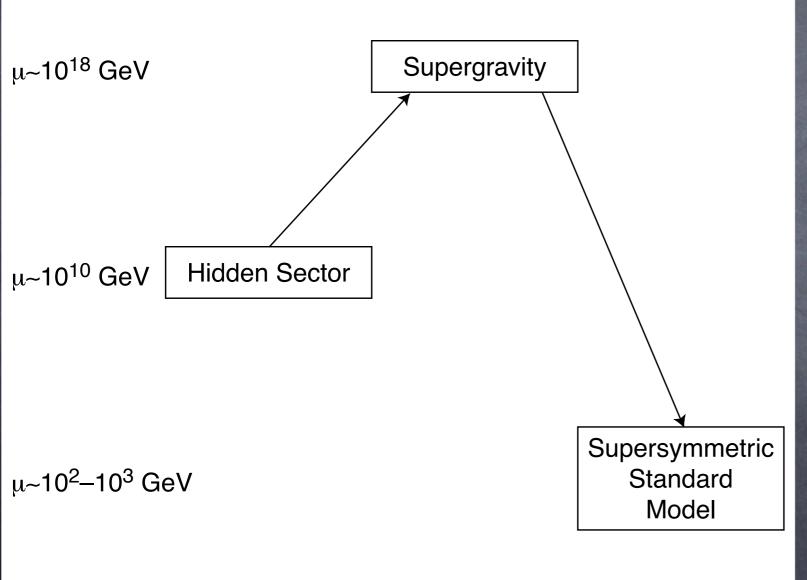
$$(5.0 \pm 0.5) \times 10^{-10} \text{ (Burles, Nollett, Turner)}$$

$$MMAP \qquad (6.5^{+0.4}_{-0.3}) \times 10^{-10}$$



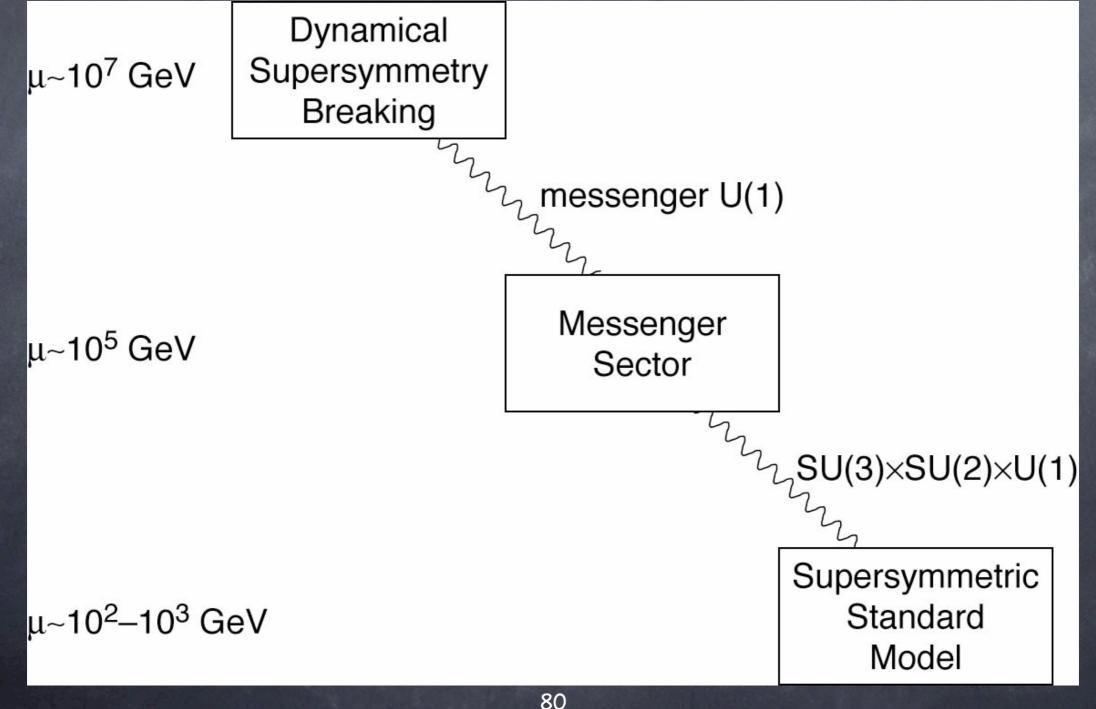
Some Information

Gravity Mediation

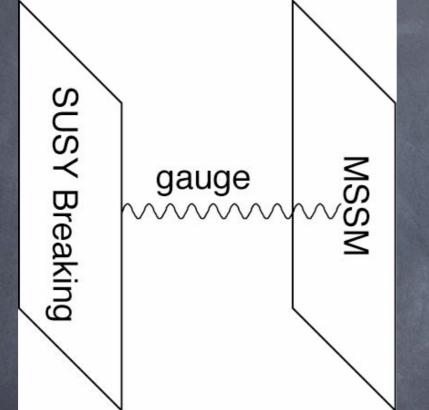


 Not really gravity, but rather gravityscale mediation

Gauge Mediation (GMSB)



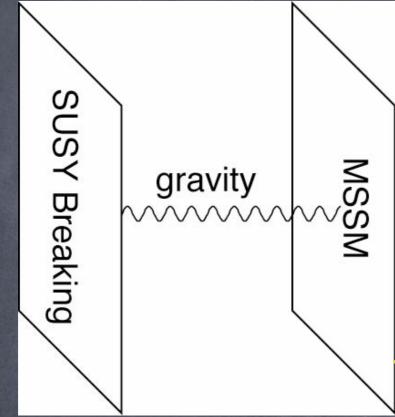
Gaugino Mediation (XMSB)



DSB in another brane
Gauge multiplet in the bulk

- Gauge multiplet learns
 SUSY breaking first,
 obtains gaugino mass
- MSSM at the compactification scale with gaugino mass only
 Scalar masses generated by RGE
 Kaplan, Kribs, Schmaltz
 Chacko, Luty, Nelson, Ponton

Anomaly Mediation (AMSB)



Randall, Sundrum Giudice, Luty, HM, Rattazzi no direct coupling between two sectors
Supersymmetry breaking in the chiral compensator <S>=1 +θ²m_{3/2}
d⁴θSS̄φ^{*}φ + ∫ d² (S³λ_{ijk}φ_iφ_jφ_k + ¹/_{g²}W_αW^α)
can be scaled away φ→φ/S
but the UV cutoff acquires

S: $\Lambda_{UV} \rightarrow \Lambda_{UV}S$

 SUSY breaking through cutoff dependence: superconformal anomaly

SUSY spectra

