

# **Stability**

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**with(out) Gauss-Bonnet  
Raidion stabilizac<sup>i</sup>on  
interactions and infiltration**

- ★ radiion - a scalar field related to that distance  $\rightarrow$  to be stabilized
- ★ distance between the branes should be stabilized
- ★ many 5d models with 4d branes have been discussed since then
- ★ Horava & Witten: 11d model with 10d branes (M-theory motivated)
- ★ basic idea: we live on a brane in a higher-dimensional space-time

## motivation

$$\phi(y) = \Phi \quad (*)$$

$$ds^2 = a(y)^2 \left\{ -dt^2 + e^{2Ht} g_{ij} dx_i dx_j + dy^2 \right\} \quad (*)$$

★ ansatz for the metric and the scalar field

$$\left\{ (\Phi) U^i (y - y_i) \sum_{j=1}^i - (\Phi) \Lambda - \frac{1}{2} (\Delta \Phi)^2 - V(\Phi) \right\} \frac{1}{2k^2} [R + \alpha R_{GB}^2] = S$$

← 5d model described by the action ( $S_1/\mathbb{Z}^2$  orbifold)  
(Goldberger & Wise)

★ modeling the radion by introducing an additional bulk scalar field

$$R_{GB}^2 = R^2 - 4R_{uv}R_{uv} + R_{uv\rho\sigma}R_{uv\rho\sigma}$$

← Gauss-Bonnet (GB) term  
(\*)  $\alpha'$  expansion in string theories

★ interactions of the higher order in the curvature tensor

extension considered here

$$\star \text{ tensor bc.} \leftarrow \lim_{y \rightarrow y_i^\pm} \frac{a}{a'} = \pm \frac{1}{6} U^i$$

$$\star \text{ scalar bc.} \leftarrow \lim_{y \rightarrow y_i^\pm} \frac{a}{\phi'} = \pm \frac{2}{1} U^i$$

$$\circledast \text{ where } \xi = a^2 - 4a \left( \frac{a}{a'} \right)^2$$

$$\star \text{ tensor eom.} \leftarrow \left\{ \frac{a''}{a'} - 2 \left( \frac{a}{a'} \right)^2 + H^2 \right\} \left[ 1 + \frac{a}{\xi} \right] - \frac{2}{1} \phi'^2 + a^2 V = 0$$

$$\star \text{ scalar eom.} \leftarrow \left\{ \frac{a''}{a'} - 2 \left( \frac{a}{a'} \right)^2 + H^2 \right\} \frac{\xi}{a^2} + \frac{3}{1} \phi'^2 = 0$$

$$\star \text{ scalar eom.} \leftarrow \phi'' + 3 \frac{a}{a'} \phi' - a^2 V' = 0$$

background equations of motion & boundary conditions

## scalar perturbations

★ generalized longitudinal gauge

$$\circledast \quad ds^2 = a^2 \left\{ (1 + 2F_1) \left[ -dt^2 + e^{2Ht} g_{ij} dx_i dx_j \right] + (1 + 2F_2) dy^2 \right\}$$

$$\circledast \quad \Phi = \phi + F_3$$

★ linearized Einstein equations

$$\circledast \quad \zeta F_1' + \frac{a}{a'} F_2 = 0$$

$$\circledast \quad (\zeta F_1)' + \frac{3}{1} a^2 \phi' F_3 = 0$$

$$0 = F_3 \left\{ \phi \frac{a}{a'} + \frac{3}{1} \phi'' F_1 \right\} + \frac{a^2}{\zeta} \left\{ (\square + 4H^2) F_1 + 4 \frac{a}{a'} F_1' - 4 \left( \frac{a}{a'} \right)^2 F_2 + \right.$$

★ boundary conditions

$$\circledast \quad \lim_{y \rightarrow y_i^+} \left\{ F_1^3 - F_2 \phi' \right\} = \mp \frac{1}{2} a F_3 U''$$

$$\left\{ \frac{\phi_{''}}{\phi} \pm \frac{a}{a_{''}} \right\} \frac{1}{2} a U_{''}^{1/2} \mp \frac{y \rightarrow y_1^+ / y_2^-}{\lim_{y \rightarrow y_1^+ (y_2^-)} F_{m^2}} = \textcolor{blue}{\text{where } b_{1/2} =}$$

$$\pm b_1^{(2)} \lim_{y \rightarrow y_1^+ (y_2^-)} F_{m^2} + \xi' \left\{ F_{m^2} + m^2 + 4H^2 \right\} = 0$$

★ eliminating  $F_2, F_3$  (and  $F_1'$ )  $\leftarrow$  boundary conditions

$\leftarrow$  scalars mass squared in the effective 4d description

★ separation constant  $m^2$

$$0 = \left\{ \frac{\xi''}{\xi'} - \frac{\xi a'}{\xi'} - 2 \frac{\xi \phi_{''}}{\xi' \phi'} - \frac{a^3 \xi'}{a' \xi'} - 3 a' \xi^2 (\phi')^2 + m^2 + 4H^2 \right\} F_{m^2} +$$

★ dynamical equation of motion  $\leftarrow F_{m^2}'' + 2 \left\{ 2 \frac{\xi}{\xi'} - \frac{a}{a'} - 2 \frac{\phi_{''}}{\phi'} \right\} F_{m^2}$

★ defining  $F_1(t, \vec{x}, y) = \sum_{m^2} F_{m^2}(y) \int d^3 k f^{(m^2, k)}(t) e^{i k \vec{x}}$

★ perturbations are not independent -  $F_2$  and  $F_3$  can be eliminated

variables elimination and separation

$$0 = (\gamma_i \partial_y^i - \partial_y(\gamma_i)) \frac{\partial u}{\partial \partial_y^i}$$

★ with non-standard boundary conditions

$$\chi = m^2 + 4H^2 \quad (*)$$

$$b = \frac{2a'\zeta^2}{a^2\zeta'} \quad (*)$$

$$d = \frac{2a\phi'^2}{3} \quad (*)$$

i.e. Sturm-Liouville differential equation, where

$$d\chi = \partial^b + (\partial^d) -$$

★ dynamical equation becomes

$$\star \text{defining } \partial^m = \zeta F^{m2}$$

summing-up the problem



stability conditions

★ static branes ( $H = 0$ ):  $\alpha = m^2 \leftarrow$  stability for  $\alpha_0 < 0$

★ brane system is stable if

$$q_i < 0 \quad (*)$$

$$0 < \frac{(\dot{\eta})_{,i}}{(\dot{\eta})_{,j} \xi_{,i}} \quad (*)$$

$$\phi'(\eta) \neq 0 \quad (*)$$

← sufficient & necessary conditions

## role of Gauss-Bonnet interactions

★ stability conditions  $\rightarrow$  addition of GB interactions unimportant?

★ numerics  $\rightarrow$  solutions with small  $\alpha \neq 0$  differ from those with  $\alpha = 0$

★ qualitative analysis

⊗ GB with  $\alpha < 0$ : model dependent, in general worse stability

⊗ GB with  $\alpha > 0$  (as predicted by the string theory):

inter-brane distance decreases, radion mass squared increases  
 $\leftarrow$  stability of the brane positions improves!

★ quantitative analysis: numerics