String Theory and Particle Physics

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Hamburg, April 2003

Introduction: The Standard Model

Particle physics is based on an experimentally extremely successful theory

the Standard Model (SM)

Properties of the Standard Model:

Spectrum:

• matter fields: (spin = 1/2)

3 families of quarks and leptons with measured masses and couplings,

• gauge bosons: (spin = 1)

gluons, W^{\pm}, Z^0, γ of gauge group

$$G_{SM} = SU(3) \times SU(2) \times U(1)_Y \xrightarrow{\langle H \rangle} SU(3) \times U(1)_{em}$$

• Higgs field H (spin = 0)

➡ experimental situation:

- good qualitative/quantitative agreement in SU(3) sector
- impressive quantitative agreement in $SU(2) \times U(1)$ sector (electroweak precision data)
- no direct observation of Higgs boson yet
- observation of neutrino masses (first modification of SM)

➡ mathematical framework:

local quantum field theory (QFT) with spontaneously broken non-Abelian gauge symmetry

- electroweak precision data test 'rules' of QFT
- gauge symmetry ensures consistency of quantization and fixes/constrains the couplings

SM leaves a number of open questions:

- what determines G_{SM} and the spectrum of particles ?
- many (19) free parameters: masses, couplings, ...
- what sets the weak scale ?
- what keeps it stable ?
- gravity cannot be consistently turned on (general relativity is not a renormalizable quantum field theory)

\Rightarrow belief:

Standard Model is only an 'effective' theory above some scale $l \leq 10^{-18} m$

Below this scale: new phenomena and new theory.

The theory of gravity: General Relativity

GR is the field theory of the gravitational interaction

- the gravitational field is the metric tensor $g_{\mu\nu}$ of space-time
- energy and momentum act as sources of the gravitational field
- Einstein's field equations

$$R_{\mu\nu} - \frac{1}{2}g_{\mu\nu}R - \Lambda g_{\mu\nu} = 8\pi G c^{-4} T_{\mu\nu}$$

• measured constants:

G: Newton's constant , Λ : cosmological constant

GR successfully governs cosmology and astrophysics

 \Rightarrow Standard Model of Cosmology

GR leaves a number of open questions:

- what sets the value of G why is gravitational interaction so weak compared to strong & electro-weak interactions?
- what sets the value of Λ why is it so small ?
- physics of singularities (Big Bang and black holes) ?
- quantum theory of gravity ?
- \Rightarrow belief: GR is not a fundamental theory
- \Rightarrow Theories "beyond" SM and GR:
 - Technicolor
 - Supersymmetry/Supergravity
 - Grand Unified Theories
 - String Theory

Supersymmetry [Wess, Zumino]

- Symmetry among fermions and bosons;
- ➡ generalization of Poincaré-algebra

 $\{Q^{I}, Q^{\dagger J}\} = \gamma^{\mu} p_{\mu} \delta^{IJ}, \quad I, J = 1, \dots, N,$

Q is generator of supersymmetry transformation.

- ➡ Properties:
 - fermions & bosons sit in the same multiplet
 ⇒ enlargement of particle spectrum;
 - quantum corrections are "tamed"
 - \Rightarrow light Higgs boson is 'natural' and predicted,
 - \Rightarrow strongly coupled QFT can be better controlled.
 - consistent with electro-weak precision data;
 - dark matter candidate: LSP
 - gauge coupling unification

(super-) GUTs:

[Georgi, Glashow]

8

Generalizes the gauge principle of the SM:

• all interactions are unified in one gauge group

 $G_{\rm GUT} \supset G_{\rm SM} = SU(3) \otimes SU(2) \otimes U(1) ,$

(e.g.
$$G_{\text{GUT}} = SU(5), SO(10), E_6$$
)

• quarks and leptons are in the same multiplet

necessary condition: unification of gauge couplings

 \Rightarrow supersymmetric GUTs at

$$M_{GUT} \sim 3 \cdot 10^{16} GeV$$

- ➡ Properties:
 - predicts decay of the proton \Rightarrow instability of matter
 - suggests light neutrinos

$$m_{\nu} \sim \frac{M_{Z^0}^2}{M_{GUT}} \sim \mathcal{O}(10^{-3} \text{eV})$$

Supergravity

Supergravity \equiv gauged (local) supersymmetry

 \Rightarrow necessary to introduce fermionic gauge field, the gravitino (s=3/2) forms together with graviton a supermultiplet

 \Rightarrow local supersymmetry only exists when gravity is turned on

- \Rightarrow If gravitino (& Higgs) exist all spins $0 \le s \le 2$ occur
- supersymmetric SM with soft supersymmetry breaking can be embedded into spontaneously broken supergravity
- Supergravity not settled but not expected

10

At what scale is a quantum gravity necessary?

$$\lambda_{\rm Compton} = \frac{\hbar}{Mc} \approx R_{Schwarz} = \frac{MG}{c^2}$$

Planck mass:MPlanck energy:EPlanck length: l_1 Planck time: t_1

$$M_{PL} = \sqrt{\frac{\hbar c}{G}} \approx 10^{-5}g$$
$$E_{PL} = c^2 M_{PL} \approx 10^{19} GeV$$
$$l_{PL} = \sqrt{\frac{G\hbar}{c^3}} \approx 10^{-35}m$$
$$t_{PL} = l_{PL}/c \approx 5 \cdot 10^{-44}s$$

$relevant \ in$

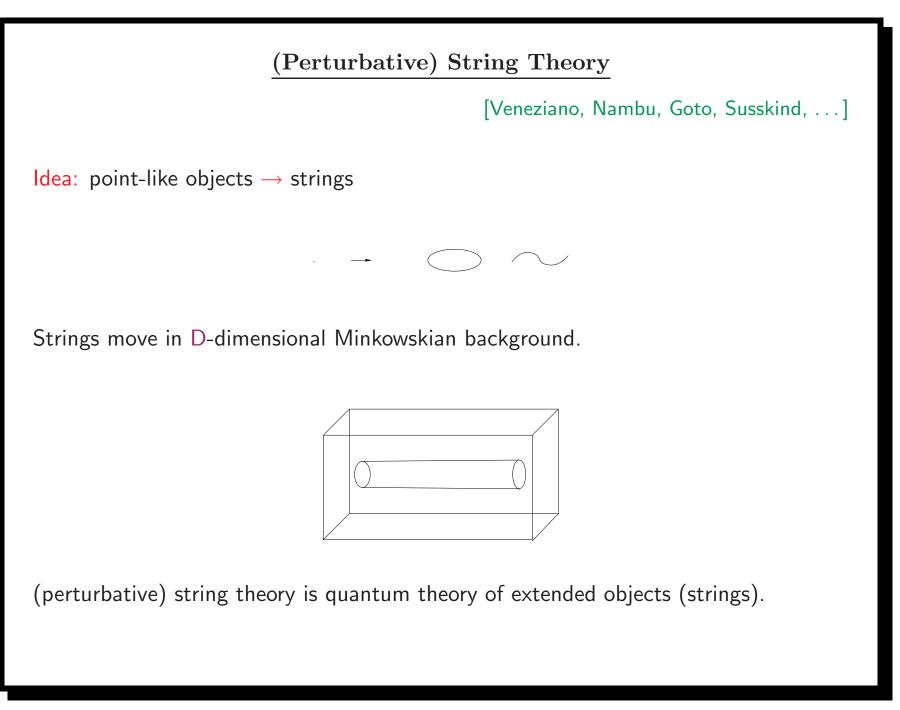
- history of the (very) early universe $\approx 10^{-43}$ sec after big bang
- physics of black holes of mass $M \approx M_{PL}$

conjecture:

at length scales

$$l \approx l_{PL} = \sqrt{\frac{\hbar G}{c^3}} \approx 10^{-35} m$$

a completely new concept is necessary to describe nature.



Quantum excitations:

• finitely many massless excitations L:

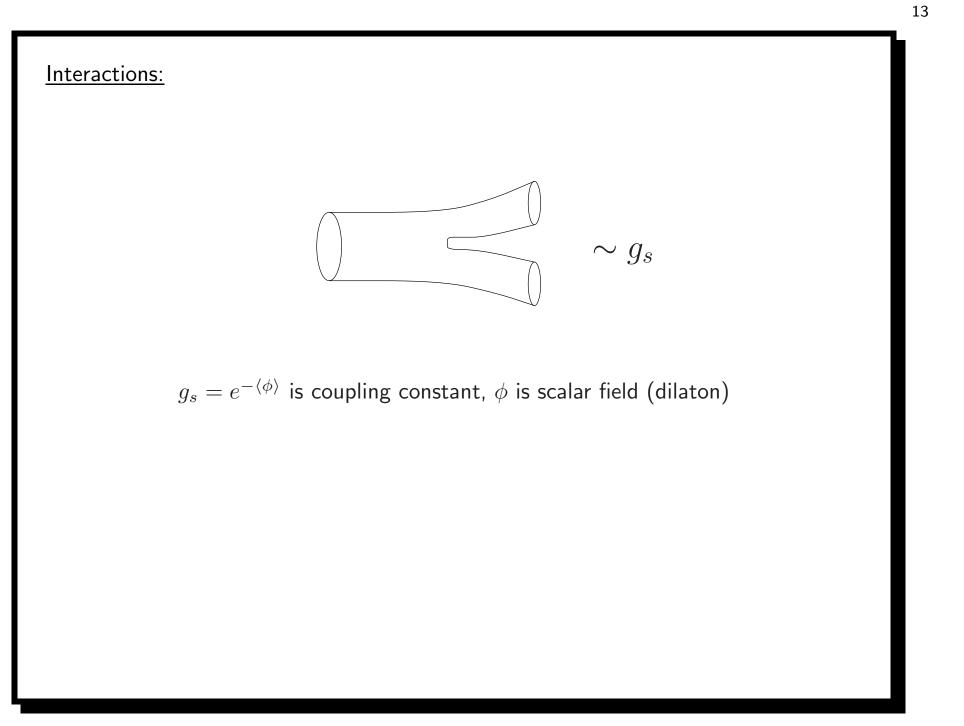
 $\begin{array}{lll} \mathsf{s}=2 & \to & \mathsf{graviton} \\ \mathsf{s}=3/2 & \to & \mathsf{gravitino} \\ \mathsf{s}=1 & \to & \mathsf{gauge\ bosons} \end{array}$

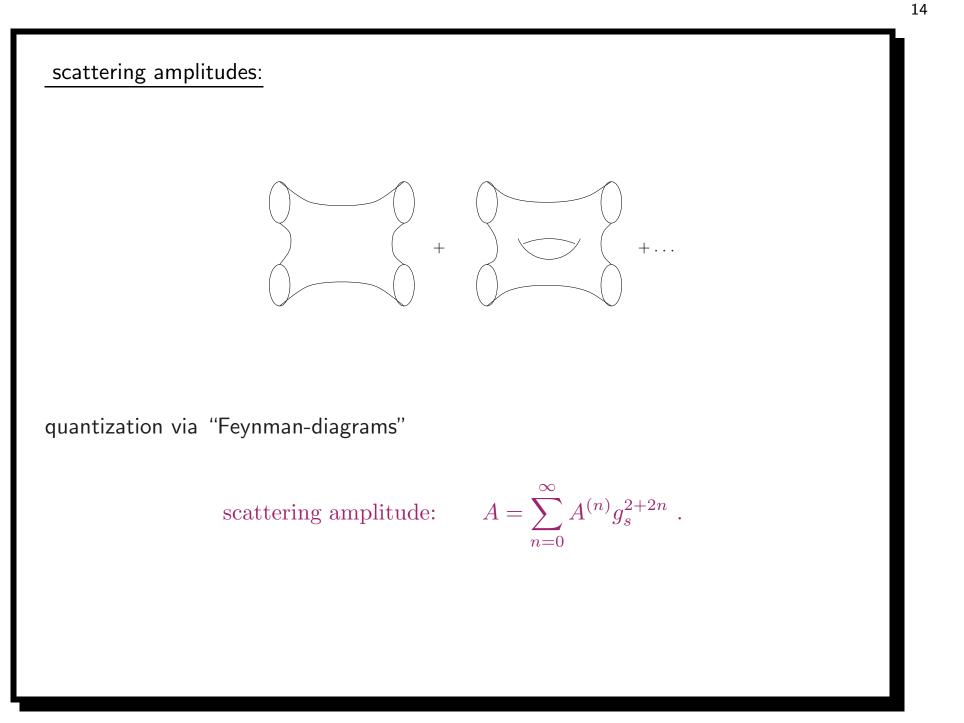
- $s = 1/2 \rightarrow$ fermions (quarks & leptons)
- $s = 0 \longrightarrow Higgs, ...$
- $\bullet\,$ infinitely many massive excitations H

$M \sim n \cdot M_S$

 $M_{\rm S}$ = characteristic scale of string theory (tension of the string)

 \Rightarrow soft UV behavior





Results:

- spectrum contains non-Abelian gauge theory with families of chiral fermions coupled to gravity
- r for scattering processes with $p \ll M_S$:

string theory $\xrightarrow{p \ll M_S}$ QFT & GR $A_{string} \longrightarrow A_{QFT,GR}$ (with $M_S \sim M_{Pl}$)

 \Rightarrow QFT & GR are low energy limit of string theory.

- ⇒ g_s is free parameter and one can choose $g_s \ll 1$ ⇒ perturbative evaluation of A possible
- $rac{1}{2}$ amplitudes $A^{(n)}$ are UV-finite ?

⇒ string theory is candidate for perturbative quantum gravity coupled to non-Abelian gauge theory

Properties

- Spectrum is supersymmetric
 - \Rightarrow necessary for consistency ?
- ➡ unitarity of scattering amplitudes

 $D \leq 10$

- D = 10: 5 different string theories: IIA, IIB, I, Het. $E_8 \times E_8$, Het. SO(32)
- *D* < 10: families of theories geometrical compactification:

 $\mathcal{M}^{(10)} = \mathcal{M}^{(D)} \otimes K^{(10-D)}$

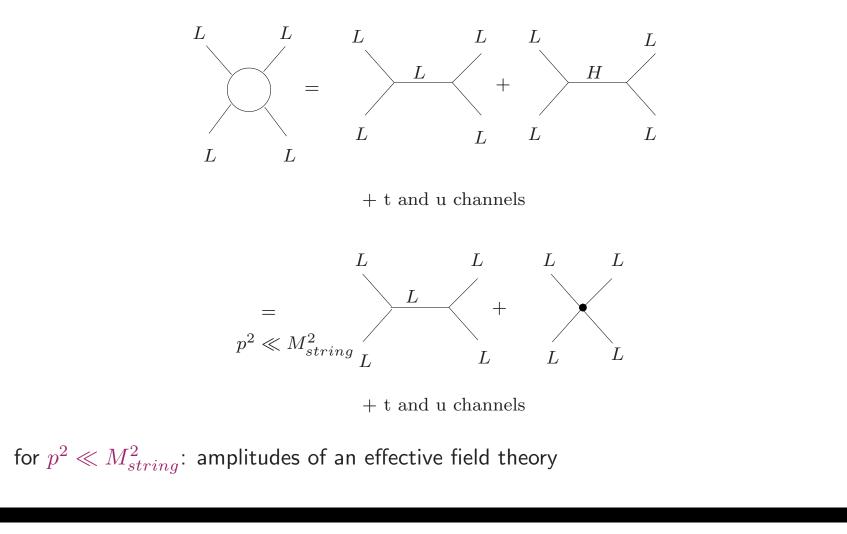
 \Rightarrow

consistency demands K = Calabi-Yau manifold holonomy group of $K \Leftrightarrow$ number of supersymmetries

 $\Rightarrow \text{ what chooses } D ?$ $\Rightarrow \text{ what chooses } K ?$

Contact with SM: The effective action

Integrating out the heavy modes H results in the low energy effective action $\mathcal{L}_{\rm eff}(L)$ of the light modes L



Problems

- Supersymmetry unbroken
- \clubsuit quarks, leptons and Higgs massless, $SU(2)\times U(1)$ unbroken
 - \Rightarrow what generates small masses ?
 - \Rightarrow what generates the hierarchy $\frac{m_Z}{m_{Pl}}~\approx 10^{-17}$
- $r \Rightarrow$ effective potential V has flat directions (moduli T)
 - \Rightarrow continuous vacuum degeneracy parameterized by $\langle T \rangle$
- r Yukawa Y and gauge couplings g dynamically determined

 $Y = Y(\langle T \rangle)$, $g = g(\langle T \rangle)$

but free parameters.

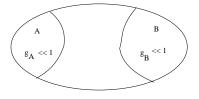
hope: cured by non-perturbative effects.

Non-perturbative Aspects of String Theory

[Hull, Townsend, Witten, ...]

conjecture:

different string theories are dual description of one quantum theory:



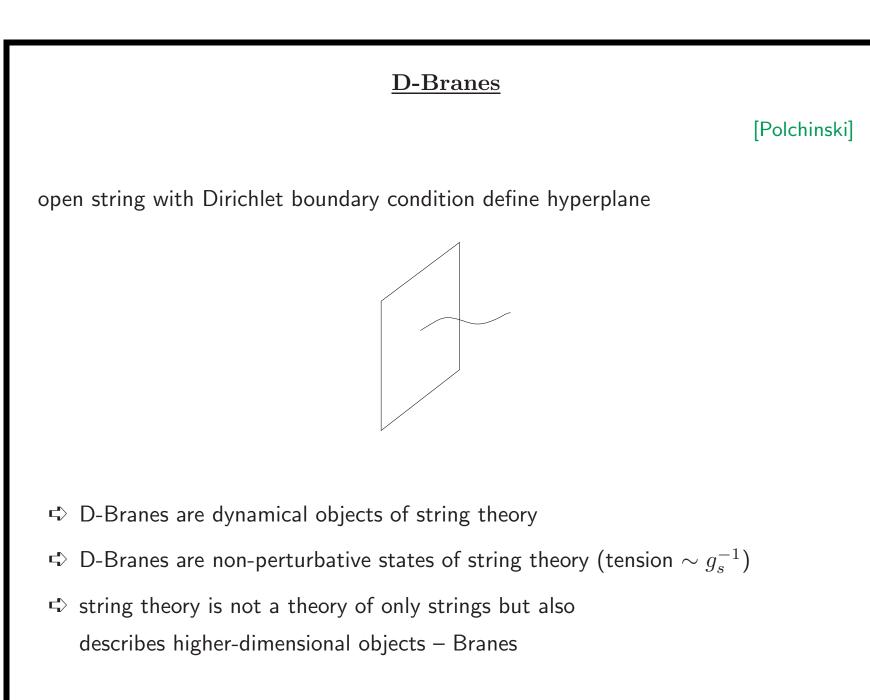
perturbative spectrum A \Leftrightarrow non- perturbative spectrum B perturbative spectrum B \Leftrightarrow non- perturbative spectrum A

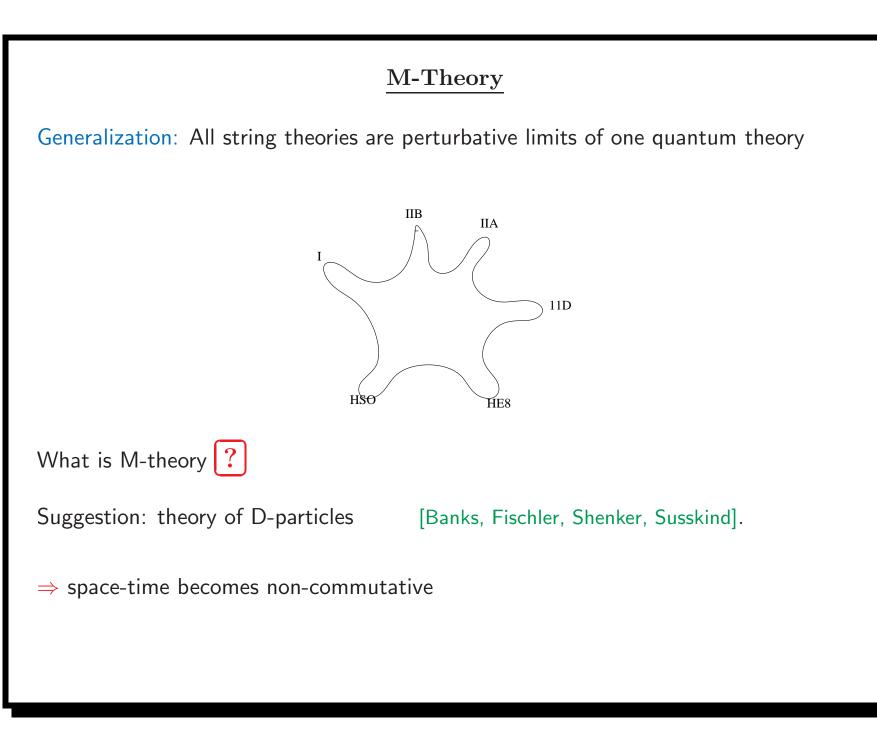
difficult to prove but successful checks on 'protected' couplings

(such couplings do exist in supersymmetric theories)

Non-perturbative states of string theory: D-branes







Perspectives – where do we go from here?

- ➡ string theory as a fundamental theory
 - conceptual
 - Particle Physics
 - GR/cosmology
- String theory as a technical tool
 - QFT
 - Mathematics

string theory as a fundamental theory

- conceptual
 - define and understand M-theory $\ensuremath{\mathsf{or}}$
 - $-\,$ define and understand non-perturbative string theory
- Particle Physics
 - study compactifications of string theory and their non-perturbative properties
 - is our universe a D-brane? \Rightarrow study Brane-World scenarios
- GR/Cosmology
 - study time-dependent (cosmological) string backgrounds
 - develop string scenarios for the Big Bang
 - embed de-Sitter backgrounds in string theory
 - study quantum Black Hole physics

string theory as a technical tool

 \triangleleft QFT

• use string theory to organize the Feynman perturbation theory

[Bern, Dixon, Kosower]

- use string theory as a regulator \Rightarrow lessons for supersymmetric QFT
- learn about strongly coupled (supersymmetric) $QFTs \Rightarrow (s)QCD$
 - AdS/CFT correspondence (N = 4)
 - Seiberg/Witten (N = 2)
 - Dijkgraaf/Vafa (N=1)

and develop new tools

String theory and Mathematics

- point-like particle ≡ probe of continuous space-time geometry
 ⇒ relation with Riemannian geometry
- string as probe: sees coarser structure
 ⇒ development of quantum geometry

surprising results:

- mirror symmetry in Calabi-Yau manifolds
- computation of number of holomorphic curves on Calabi-Yau manifolds
- development of quantum cohomology

[Einstein,Hilbert]

[Kontsevich, Manin, ...]

Summary

- string theory unifies all interactions and provides perturbative quantum gravity.
- qualitative agreement with generalizations of the Standard Model but no quantitative agreement yet.
- biggest problems:
 - hierarchical supersymmetry breaking,
 - determination of the ground state,
 - cosmological constant.
- non-perturbative properties at least partially under control.
- phenomenological implications are being investigated.