THÉORIE

DU

MOUVEMENT DE LA LUNE,

PAR CH. DELAUNAY, MEMBRE DE L'INSTITUT IMPÉRIAL DE FRANCE.

TOME PREMIER.

PARIS,

MALLET-BACHELIER, IMPRIMEUR-LIBRAIRE

DES COMPTES RENDUS HEBDOMADAIRES DES SÉANCES DE L'ACADÉMIE DES SCIENCES,

QUAL DES GRANDS-AUGUSTINS, 55.

1860.

 $\Delta = Z$

.

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CHAPITRE V.

DÉTAIL DES 57 OPÉRATIONS EFFECTUÉES POUR FAIRE DISPARAITRE LES TERMES LES PLUS IMPORTANTS DE LA FONCTION PERTURBATRICE.

Nous avons expliqué, dans le chapitre III, la marche que nous avons suivie pour faire disparaître successivement de la fonction perturbatrice les divers termes capables de fournir des inégalités d'un ordre inférieur au quatrième. Nous avons dit qu'il nous a fallu, pour cela, effectuer 57 opérations distinctes, dont chacune se traduit en définitive par un changement de variables. Nous nous proposons de donner ici le détail des calculs qui se rapportent à ces diverses opérations, pour chacune desquelles nous n'avons autre chose à faire que d'appliquer une des quatre règles données aux n^{es} 29, 30 et 31.

On se rappelle que les équations différentielles qu'il s'agit d'intégrer sont les suivantes :

$\frac{d\mathbf{L}}{dt}=\frac{d\mathbf{R}}{dt},$	$\frac{d\mathbf{G}}{dt} = \frac{d\mathbf{R}}{dg},$	$\frac{d\mathbf{H}}{dt} = \frac{d\mathbf{R}}{dh},$
$\frac{dl}{dt} = -\frac{d\mathbf{R}}{d\mathbf{L}},$	$\frac{dg}{dt} = -\frac{d\mathbf{R}}{d\mathbf{G}};$	$\frac{dh}{dt} = -\frac{d\mathbf{R}}{d\mathbf{H}}.$

La fonction perturbatrice R qui entre dans ces équations, est donnée en fonction du temps t et des éléments variables de la Lune par le développement du n° 14; mais on peut aussi la prendre dans le chapitre IV, en ayant soin de ne conserver, dans les coefficients des divers termes, que les parties qui*ne sont pas accompagnées d'indications en petits chiffres placés au dessous.

Des six variables L, G, H, l, g, h, auxquelles se rapportent les équations différentielles précédentes, il n'y a que les trois dernières qui entrent explicitement dans la valeur de R. Les trois autres L, G, H, y sont remplacées par les éléments a, e, γ , auxquels elles sont liées par les relations

L =
$$\sqrt{a_{2'}}$$
, G = L $\sqrt{1 - e^2}$, II = G $(1 - 2\gamma^2)$.
T. I. 33

	n		
2	b	τ	٠

NUMÉROS	ARGUMENTS DIS TERMES	ORDBES
des	que ces opérations sont destinées à faire	deces
opérations.	disparaître de R.	opérations.
	•	
38	3/	3° ordre.
39	2h + 2g + 5l - 2h' - 2g' - 2l'	3° ordre.
40	2h + 2g - l - 2h' - 2g' - 2l'	3° ordre.
41	2h + 2g - 2h' - 2g' - 2l	1" ordre.
42	ah + ag - ah' - ag' - 3l'	2º ordre.
43	2h + 2g - 2h' - 2g' - l'	2° ordre.
44	2h + 2g - 2h' - 2g' - 4l'	3º ordre.
45	2h + 2g - 2h' - 2g'	3° ordre.
4G	h+g-h'-g'-l'	2* ordre.
47	h+g-h'-g'-at	3° ordre.
48	h+g-h'-g'	2º ordre.
49	2g	a* ordre.
50	2h + 2g - 2h' - 2g' - 2l' (une seconde fois)	3° ordre.
51	2g (une seconde fois)	3 ^e ordre.
52	2h - 2h' - 2g' - 2l'	I" ordre.
53	2h - 2h' - 2g' - 3l'	2* ordre.
54	2h - 2h' - 2g' - l	'a* ordre.
55	2h - 2h' - 2g' - 4l'	3° ordre.
56	2h - 2h' - 2g'	3° ordre.
57	l', al' , $3l'$, $4l'$ (une seconde fois)	3° ordre.

I" OPÉRATION

destinée à faire disparaître les termes (2), (3), (4), (5) et (6) de R.

Prenons dans R les termes (2), (3), (4), (5) et (6) *, dans lesquels les arguments sont l', 2l', 3l', 4l', 5l', et supposons que R se réduise à ces termes seuls, de sorte que l'on ait

$$\begin{aligned} \mathbf{R} &= m^{2} \frac{a^{3}}{a^{2}} \left[\frac{3}{4} e^{\prime} - \frac{9}{2} \gamma^{3} e^{\prime} + \frac{9}{8} e^{2} e^{\prime} + \frac{27}{32} e^{\prime \prime} + \frac{9}{2} \gamma^{4} e^{\prime} - \frac{27}{4} \gamma^{2} e^{2} e^{\prime} - \frac{81}{16} \gamma^{2} e^{\prime 3} + \frac{81}{64} e^{2} e^{3} + \frac{261}{256} e^{\prime 3} \\ &+ \frac{27}{4} \gamma^{4} e^{2} e^{\prime} + \left(\frac{45}{64} e^{\prime} - \frac{225}{16} \gamma^{2} e^{\prime} + \frac{225}{64} e^{2} e^{\prime} \right) \frac{a^{2}}{a^{2}} \right] \cos t \end{aligned}$$

De ces valeurs de L, G, H, on déduit $\frac{da}{dL} = \frac{1}{an} \left\{ 2 + \left(\frac{359}{8} - \frac{375}{2}\gamma^2 + \frac{45371}{128}e^2 + \frac{5385}{16}e^2\right) \frac{n^6}{n^4} + \left(140 - 396\gamma^2 + \frac{18411}{8}e^2 + 1890e^2\right) \frac{n^6}{n^4} + \frac{18411}{8}e^2 + 1890e^2\right) \frac{n^6}{n^4} + \frac{18411}{8}e^2 + 1890e^2$ $+\frac{41875}{65}\frac{\pi^{\prime\prime}}{\mu^{\prime\prime}}+\frac{483281}{288}\frac{\pi^{\prime\prime}}{\mu^{\prime\prime}}$ $\frac{da}{dG} = -\frac{1}{au} \left\{ \left(\frac{387}{4} - \frac{1473}{4} \gamma^2 - \frac{22933}{128} e^2 + \frac{5805}{8} e^2 \right) \frac{u^n}{n^4} + \left(420 - 1476\gamma^2 - \frac{5997}{8} e^2 + 5670 e^2 \right) \frac{u^n}{n^4} \right\}$ $+ \frac{126279}{64} \frac{n^4}{n^6} + \frac{1880475}{288} \frac{n^6}{n^7} \Big\langle ,$ $\frac{da}{dH} = -\frac{1}{an} \left\{ \left(\frac{33}{4} - \frac{69}{4} \gamma^2 - \frac{771}{8} e^2 + \frac{495}{8} e^{i2} \right) \frac{n^{i_1}}{n^*} + (40 - 120 \gamma^2 - 399 e^2 + 540 e^{i2}) \frac{n^{i_1}}{n^*} \right\}$ $= \frac{\pi}{n^2} \cdot \frac{10425}{128} e^{\prime 2} \cos 3\ell,$ $+\frac{6191}{32}\frac{n^{n}}{n^{4}}+\frac{94481}{144}\frac{n^{2}}{n^{2}}\},$ $\frac{de}{d1} = \frac{1}{n^2 ne} \left\{ 1 - e^2 + \left(\frac{927}{32} - \frac{691}{8}\gamma^2 - \frac{36549}{128}e^2 + \frac{13955}{64}e^{\gamma^2}\right) \frac{n^4}{n^4} + \frac{412}{3}\frac{n^5}{n^2} + \frac{496405}{768}\frac{n^{44}}{n^6} \right\},$ $\frac{de}{dG} = -\frac{1}{a^2 ne} \left\{ 1 - \frac{1}{2}e^2 - \frac{1}{8}e^2 - \frac{1}{16}e^4 \right\}$ $+\left(\frac{927}{3a}-\frac{691}{8}\gamma^2-\frac{57693}{188}e^3+\frac{13905}{64}e^5\right)\frac{n''}{n'}+\frac{412}{3}\frac{n''}{n'}+\frac{496405}{768}\frac{n''}{n''}\Big\},$ $\frac{de}{dH} = \frac{1}{a^{4}ne} \cdot \frac{757}{32} e^{3} \frac{n^{4}}{n^{4}},$ $\frac{d\gamma}{dL} = -\frac{1}{n^2 n \gamma} \cdot \frac{17}{4} \gamma^2 \frac{n^2}{n^4},$ $\frac{d\gamma}{dG} = \frac{1}{4a^2 n\gamma} \left\{ 1 - 2\gamma^2 + \frac{1}{2}e^2 - \gamma^2 e^2 + \frac{3}{8}e^4 - \frac{3}{4}\gamma^2 e^4 + \frac{5}{16}e^4 \right\}$ $+\left(\frac{43}{16}+\frac{879}{8}\gamma^2-\frac{1427}{64}e^2+\frac{645}{32}e^{r^2}\right)\frac{n^{r_1}}{n^4}+\frac{61}{6}\frac{n^{r_2}}{n^5}+\frac{8021}{102}\frac{n^{r_4}}{n^6}\Big\},$

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CHAPITRE V. -- 10" OPERATION.

 $\frac{d\gamma}{dH} = -\frac{1}{4n^2n^2} \left\{ 1 + \frac{1}{2}e^2 + \frac{3}{8}e^4 + \frac{5}{16}e^6 + \left(\frac{43}{16} - \frac{127}{4}\gamma^2 - \frac{1427}{64}e^3 + \frac{645}{32}e^6\right) \frac{n^6}{n^4} + \frac{61}{6}\frac{n^6}{n^5} + \frac{8021}{102}\frac{n^6}{n^5} \right\}$

CHAPITRE V. - 57" OPÉRATION. 875

Quant aux valeurs de l, h + g + l et h, elles seront fournies par les équations différentielles

$$\frac{dt}{dt} = -\frac{d\mathbf{R}}{d\mathbf{L}}, \qquad \frac{d(\mathbf{h} + \mathbf{g} + t)}{dt} = -\frac{d\mathbf{R}}{d\mathbf{L}} - \frac{d\mathbf{R}}{d\mathbf{G}} - \frac{d\mathbf{R}}{d\mathbf{H}}, \qquad \frac{d\mathbf{h}}{dt} = -\frac{d\mathbf{R}}{d\mathbf{H}},$$

qui, en vertu des valeurs de $\frac{da}{dL}$, $\frac{da}{dG}$, $\frac{da}{dH}$, $\frac{de}{dL}$, ..., deviennent

$$\frac{d'}{d'} = -\frac{u^2}{n'} \left[-\frac{225}{2} \gamma^i e' + \frac{225}{4} \gamma^i e' e' + \left(\frac{675}{16} e' - \frac{243}{2} \gamma^j e' + \frac{2025}{32} e^2 e' + \frac{8475}{128} e'^i \right) \frac{n'}{n'} \right]$$
$$+ \left(\frac{9561}{64} e' - \frac{19197}{32} \gamma^2 e' + \frac{63801}{128} e^2 e' \right) \frac{n'^2}{n^2} + \frac{448647}{512} e' \frac{n'^3}{n^2} + \frac{9481465}{2048} e' \frac{n''}{n'} \right] \cos \frac{n'^2}{n'} \left[\frac{825}{8} e^{i2} - 297 \gamma^2 e'^2 + \frac{2475}{16} e^2 e'^2 + \frac{38932}{128} e'^2 \frac{n'}{n} + \frac{2611785}{1024} e'^2 \frac{n'^2}{n^2} \right] \cos \frac{n'}{n'}$$

$$\frac{d(h+g+I)}{dt} = -\frac{a^{\prime 3}}{a^{\prime}} \left[\frac{8i}{4} \gamma^{\prime} r^{\prime} + \frac{2025}{16} e^{2} e^{\prime} - \frac{405}{8} \gamma^{\prime} e^{\prime} - \frac{729}{2} \gamma^{2} e^{2} e^{\prime} - \frac{6075}{128} e^{4} e^{\prime} - \left(\frac{735}{128} e^{-} - \frac{4779}{32} \gamma^{2} r^{\prime} - \frac{112509}{128} e^{2} e^{\prime} + \frac{28665}{128} e^{-3} \right) \frac{n^{\prime}}{n} - \left(\frac{3783}{16} e^{\prime} - \frac{87525}{128} \gamma^{2} e^{\prime} - \frac{2962593}{512} e^{2} e^{\prime} \right) \frac{n^{\prime 2}}{n^{2}} - \frac{4571}{4} e^{\prime} \frac{n^{\prime 3}}{n^{2}} - \frac{184357}{48} e^{\prime} \frac{n^{\prime 4}}{n^{4}} + \frac{2475}{64} e^{\prime} \frac{n^{2}}{n^{2}} \right] \cos \theta$$

$$-\frac{n}{n^2} \left[\frac{99}{2} \gamma^2 e^{\prime 2} + \frac{4375}{8} e^{\circ} e^{\prime 2} - \left(\frac{6013}{64} e^{\prime 2} - \frac{23977}{64} \gamma^2 e^{\prime 2} - \frac{409305}{256} e^{\circ} e^{\prime 2} \right) \frac{n}{n} - \frac{60203}{128} e^{\prime 2} \frac{n^2}{n^2} - \frac{273133}{96} e^{\prime 2} \frac{n^2}{n^2} \right] \cos 2t'$$

$$\frac{dh}{dt} = \frac{n^2}{n} \left[-\frac{225}{4} \gamma^2 e^2 e^t + \frac{225}{32} e^4 e^t + \left(\frac{27}{16} e^t - \frac{81}{8} \gamma^2 e^t - \frac{567}{16} e^2 e^t + \frac{339}{128} e^2\right) \frac{n^4}{n^4} + \left(\frac{531}{64} e^t - \frac{63}{4} \gamma^2 e^t - \frac{2997}{16} e^2 e^t\right) \frac{n^{42}}{n^2} + \frac{14709}{512} e^t \frac{n^{43}}{n^3} + \frac{323351}{2048} e^t \frac{n^{44}}{n^4} \right] \cos \ell^2$$

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Cette formule se continue a le pare suivante

THÉORIE

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MOUVEMENT DE LA LUNE,

PAR- CH. DELAUNAY, MEMBRE DE L'INSTITUT IMPÉRIAL DE FRANCE.

TOME SECOND.

PARIS,

GAUTHIER-VILLARS, IMPRIMEUR-LIBRAIRE DES COMPTES RENDUS HEBDOMADAIRES DES SÉANCES DE L'ACADÉMIE DES SCIENCES, SUCCESSEUR DE MALLET-BACHELIER.

QUAL DES GRANDS-AUGUSTINS, 55.

1867

Another 448 canonical transformations on 700 pages

Main results:

- Moon's longitude: 53 pages
- Moon's latitude: 52 pages

Remaining errors O(10⁻⁴)

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Analytical Lunar Ephemeris: Delaunay's Theory

ANDRÉ DEPRIT, JACQUES HENRARD, AND ARNOLD ROM Boeing Scientific Research Laboratories, Seattle, Washington (Received 1 December 1970)

Delaunay's constants have been substituted into our analytical solution of the main problem of lunar theory. The results are compared with Delaunay's reduced formulas. Corrections are proposed to four terms in the mean motion of the perigee, three in the mean motion of the node, 45 in the reduced expression for the latitude, and 49 in that for the longitude.

DELAUNAY'S THEORY

TABLE III. Corrections to Delaunay's solar terms in the longitude V.

No.	D	F	ı	ľ	#2	α	e	γ	e'	Delaunay's numerator	ALE numerator	Delaunay's denominator	ALE denominator
88 12 13 58 58 58 58 68 89 95 95 95 102 111 111 118 118 118 124 124	222222222222222222222222222222222222222	2 2 2 2 2 2	$ \begin{array}{c} 1\\ 1\\ -1\\ -1\\ -3\\ 3\\ -1\\ -1\\ -1\\ -1\\ -1 \end{array} $	-1 1 2 2 2 1 -1 -1 -1 2 2	$\begin{array}{c} 1\\ 1\\ 1\\ 4\\ 2\\ 2\\ 4\\ 2\\ 3\\ 5\\ 3\\ 5\\ 5\\ 1\\ 2\\ 3\\ 3\\ 4\\ 2\\ 3\\ 4\\ 2\\ 3\end{array}$	2	1 1 1 1 1 3 2 2 1 3 3 1 1 1 3 2 2 1 3 3 1 1 1 1 3 1 1 3 2 2 1 1 3 1 1 1 1 1 1 1 1	4 4 2 2 2 2 2 2 2	1 1 2 2 2 2 2 2 1 1 1 2 2 2 2 2 2 2 2	$\begin{array}{r} -351\\ -351\\ -357809\\ -6447\\ 57025\\ 149363\\ 1231\\ -149497\\ -61969\\ -1782049\\ -28021\\ -945\\ 5705\\ 173819\\ 365281\\ 19912163\\ 183\\ -46561\end{array}$		16	64
124 151 151 169 184 191 198 205 221 223 233 236	1222222222222244	222222222222222222222222222222222222	-1 -2 -2 -2 1 2 -1 -1	2 1 -1	342322432232264		1 2 2 2 1 2 2 1 2 1 1 2 1 1 2	2 2 2 2 2 2 2 2 2 2 2 2 2 2 4 4	2 1 1 1 1 1 1 1 1 1 1 1 1	$\begin{array}{r} -479'4330\\ 3785\\ 613\\ 5255\\ -6355\\ -13733\\ -81\\ -6197\\ 1881\\ -40795\\ -9\\ -45\\ 54129983\\ -60359\end{array}$	-73237483 3813 617 5043 -6215 -12837 -71 -7177 2021 -39339 9 52839679 -145415	8	16
237 237 239 240 255 275 314 369 372	444444 4444 1	2		$\frac{2}{2}$ -1 -2 1	454344632	1	3 3 1 1 2 2	2	2 2 1 2 1 1	161 2429 307749 140105 2056689 -467 2853 -680863 79689	201 14237 320549 136325 2056609 - 563 - 3715 - 797863 216189	2560	7680
399 399 405 418 419 421 432 462	1 1 3 3 3 3 5	-2 -2 -2	$2 \\ 1 \\ 1 \\ -2$	$\frac{1}{1}$ -1 1	2 4 3 3 5	1 1 1 1 1 1 1 1	2 2 1 1 2 2	2 2 2	1 1 1 1	145 25649 	210155 155 26795 -395 367957 25595 -15181 -2931935 1481	48 30720 768 128	144 7680 3072 1024

H. Poincaré, Bulletin Astronomique (1908)

MÉMOIRES ET OBSERVATIONS.

SUR LES PETITS DIVISEURS DANS LA THÉORIE DE LA LUNE;

PAR H. POINCARÉ.

1. Dans le développement de la théorie de la Lune, on voit s'introduire de petits diviseurs de la forme suivante :

$p_1 n_1 + p_2 n_2 + p_3 n_3 + p_4 n_4;$

les p sont des entiers, positifs ou négatifs; n_1 et n_2 sont les moyens mouvements de la Lune et du Soleil, n_3 et n_4 sont ceux du périgée et du nœud. Si l'on pose

$$\frac{n_2}{n_1}=m_1$$

Mais l'exposant de m est évidemment le même que celui de β . Il peut donc devenir négatif. Si donc on poussait assez loin les développements de Delaunay, on arriverait à des termes où mfigurerait à une puissance négative. Mais on n'y arriverait que quand on rencontrerait de très petits diviseurs analytiques, ce qui, nous l'avons dit, ne peut se produire que pour des termes d'ordre très élevé. C'est pour cette raison que cette circonstance a échappé à Delaunay. Two lessons:

- Canonical transformation are very useful for
 "Classical dynamics in systems with many coupled degrees of freedom"
- Do not worry too much about convergence



GEORG-AUGUST-UNIVERSITÄT Göttingen

Real Time Evolution of Quantum Many-Body Systems

Flow Equations and Unitary Perturbation Theory

Stefan Kehrein Institut für Theoretische Physik, Universität Göttingen

- 1. Real time evolution in classical mechanics: Canonical perturbation theory
- 2. Real time evolution in quantum mechanics: Unitary perturbation theory
- 3. Unitary perturbation theory in practice: Flow equation method
- 4. Applications (impurity models, local quenches):
 - Spin-boson model
 - Time-dependent ferromagnetic Kondo model
- 5. Applications (bulk models, global quenches):
 - Quantum quench in a Fermi liquid
 - Quantum quench for 1d fermions with dimerization
- 6. Outlook

Main collaborators:

A. Mielke (Univ. Heidelberg)A. Hackl (SAP)M. Möckel (Cambridge Univ.)N. Robinson (Oxford Univ.)



2. Real time evolution in classical mechanics: Canonical perturbation theory

"Real time evolution" with small anharmonic terms

$$H = \frac{1}{2}p^2 + \frac{1}{2}q^2 + \frac{g}{4}q^4$$

Ansatz: $q(t) = q^{(0)}(t) + g q^{(1)}(t) + O(g^2)$ (initial condition: q(t = 0) = 0) $\Rightarrow q^{(0)}(t) = c \sin t$ $\Rightarrow \ddot{q}^{(1)}(t) = -q^{(1)}(t) - c^3 \sin^3 t$ $\Rightarrow q^{(1)}(t) = \frac{c^3}{8} \left(-\frac{5}{4} \sin^3 t - \frac{3}{4} \cos^2 t \sin t + 3t \cos t \right)$ "Secular term"

Secular terms invalidate naive perturbation theory for large times!

Much better ... Canonical perturbation theory

Find canonical transformation $(q,p) \rightarrow (Q,P)$ that brings H to normal form:

$$H(q, p) = \frac{1}{2}p^{2} + \frac{1}{2}q^{2} + \frac{g}{4}q^{4}$$

$$\tilde{H}(Q, P) = H_{0} + \frac{3}{8}gH_{0}^{2} + O(g^{2}) \quad \text{with} \quad H_{0} = \frac{1}{2}P^{2} + \frac{1}{2}Q^{2}$$
vanishing Poisson bracket \Rightarrow exact solution of dynamics for Q(t), P(t)
(H_{0} = E_{0} \text{ conserved}) \qquad \text{possible (no secular terms):}
$$Q(t) = Q_{0} \sin(\omega t + \gamma_{0}) \quad \text{with} \quad \omega = 1 + \frac{3}{4}gE_{0}$$
Generating function: $F_{2}(q, P) = qP + g\left(\frac{5}{32}Pq^{3} + \frac{3}{32}P^{3}q\right)$

$$Q = \frac{\partial F_{2}}{\partial P} = q + g\left(\frac{5}{32}q^{3} + \frac{9}{32}P^{2}q\right) \quad , \quad p = \frac{\partial F_{2}}{\partial q} = \dots$$

Insert solution Q(t), P(t) and reexpress in terms of q(t), p(t):

$$q(t) = Q_0 \sin(\omega t + \gamma_0) - g \left(\frac{5}{32}Q_0^3 \sin^3(\omega t + \gamma_0) + \frac{3}{32}Q_0^3 \cos^2(\omega t + \gamma_0) \sin(\omega t + \gamma_0)\right) + O(g^2)$$

$$\uparrow$$
also calculation in first order in g



Problem of naive perturbation theory: naive expansion in coupling constant produces secular terms



canoncial perturbation theory naive perturbation theory in coupling

 \Rightarrow Same recipe for dynamics of quantum systems:

Perturbation theory based on unitary transformations instead of "naive" perturbation expansion

2. Real time evolution in quantum mechanics: Unitary perturbation theory

U: unitary transformation that diagonalizes the Hamiltonian (approximately)



"Forward-backward transformation"

3. Unitary perturbation in practice: Flow equations

Problem: How to construct U for non-integrable models with a continuum of energy scales? Wanted: Perturbative method for finding U

Stable expansion for systems with very different energy scales

- \rightarrow Energy-scale separation
- \rightarrow Scaling theory / Renormalization theory (K.G. Wilson)





Effective low-energy Hamiltonian (fixed point)

Problem:

 ${\rm H}_{\rm initial}$

Information on full Hilbert space required to construct unitary transformation for forward-backward transformation

Flow Equation Method

F. Wegner (1994)

S. K., The Flow Equation Approach to Many-Body Problems (Springer 2006)



Implementation of flow: Sequence of infinitesimal unitary transformations

One-parameter family of unitarily equivalent Hamiltonians generated by solving the differential equation ("flow equations")

$$\frac{dH}{dB} = [\eta(B), H(B)]$$

with H(B=0) the initial Hamiltonian and an anti-hermitean generator $\eta(B)$.

Canonical choice of generator (Wegner 1994):

 $H(B) = H_0(B)$ [diagonal part] + $H_{int}(B)$ [interaction part]

→ define anti-hermitean generator $\eta(B) = [H_0(B), H_{int}(B)]$

$$\Rightarrow \lim_{B \to \infty} \eta(B) = 0$$

 \rightarrow generates band-diagonal Hamiltonians H(B) with $B^{\text{-}1/2}$ = Λ_{feq}

Challenge: Generation of higher and higher order interaction terms

 \rightarrow need suitable expansion parameter (typically running coupling)

Advantages:

- RG-like analytical method
- Controlled solution in certain strong-coupling problems (e.g., Kondo model)
- Keeps all states in Hilbert space
 - \rightarrow correlation functions on all energy scales
 - \rightarrow important in non-equilibrium (real time evolution, steady states)
- Avoids secular terms in real time evolution!

4. Applications (impurity models): Spin-boson model

Paradigmatic 2-state system for dissipative quantum mechanics:

$$H = -\frac{\Delta}{2}\sigma_{x} + \sigma_{z}\sum_{k}\lambda_{k}(b_{k} + b_{k}^{\dagger}) + \sum_{k}\omega_{k}b_{k}^{\dagger}b_{k}$$
Environment degrees of
freedom \rightarrow Decoherence and
dissipation
Flow equation diagonalization:
$$\frac{d\lambda_{k}(B)}{dB} = -(\omega_{k} - \Delta)^{2}\lambda_{k}$$
Environment degrees of
freedom \rightarrow Decoherence and
dissipation
Energy scale
separation
$$\frac{d\Delta(B)}{dB} = -\Delta\sum_{k}\lambda_{k}^{2}\frac{\omega_{k} - \Delta}{\omega_{k} + \Delta} \operatorname{coth}(\beta\omega_{k}/2) \longrightarrow \begin{array}{c} \operatorname{Non-perturbative energy} \\ \operatorname{Scale}(\text{frequency}) \\ \Rightarrow H(B = \infty) = -\frac{\Delta_{r}}{2}\sigma_{x} + \sum_{k}\omega_{k}b_{k}^{\dagger}b_{k} + O(\lambda_{k}^{2}) \qquad \Delta_{r} \propto \left(\frac{\Delta}{\omega_{c}}\right)^{\alpha/1 - \alpha}$$

Time evolution with respect to $H(B=\infty)$ is trivial

- \rightarrow Where is dissipation/decoherence?
- → Transformation of observables!

Transformation of Observables

Transformation of observable O to diagonal basis ("Forward" transformation):

$$\frac{dO(B)}{dB} = [\eta(B), O(B)] \quad , \quad O(B = 0) = O$$

Ansatz for σ_z :

Decoherence in "conventional" framework:

System degrees of freedom become entangled with environment

Decoherence in flow equation framework:

Hamiltonian diagonal, therefore system observables become entangled with environment Equilibrium Dynamics (T=0)



Equilibrium: | $\tilde{\psi}_i$ > is trivially given as ground state |GS> of \tilde{H}

Spin operator in diagonal basis:

$$\sigma_z(B=\infty)=\sigma_x\sum_k\chi_k(B=\infty)\left(b_k^{\dagger}+b_k\right)$$

 \rightarrow Incoherent spin dynamics:

$$C_{zz}(t) = \frac{1}{2} \langle GS | \left\{ \sigma_z(B = \infty), e^{iH(B = \infty)t} \sigma_z(B = \infty) e^{-iH(B = \infty)t} \right\} | GS \rangle$$

Practical evaluation: Numerical evaluation of O(10³) ordinary differential equations (qualitative behavior: analytical calculation!)



Real time evolution

A. Hackl and S. K., Phys. Rev. B 78 (2008), J. Phys. C 21 (2009)

Forward-backward transformation (numerical implementation):

$$\sigma_{z}(t) = z(t)\sigma_{z} + y(t)\sigma_{y}$$

+ $i\sigma_{x}\sum_{k}\alpha_{k}(t)(b_{k} - b_{k}^{\dagger}) + \sigma_{x}\sum_{k}\beta_{k}(t)(b_{k} + b_{k}^{\dagger})$
+higher order terms

Evaluate for arbitrary initial state, e.g. here:

Initial state:
$$|\Psi_i\rangle = |\uparrow, \text{ relaxed bath}\rangle$$

Expectation values: $E_z(t) = \langle \Psi_i | \sigma_z(t) | \Psi_i \rangle = z(t)$

Shows exponential decay (possibly with oscillations) already in this order of flow equation calculation



Stable long time asymptotics (no secular terms):



Comparison with NRG data from Costi et al., PRA 68 (2003)

Applications (impurity models): Ferromagnetic Kondo model

A. Hackl, D. Roosen, S. K., W. Hofstetter, Phys. Rev. Lett. 102, 196601 (2009) A. Hackl, M. Vojta and S. K., Phys. Rev. B 80, 195117 (2009)

$$\begin{split} H_{i} &= \sum_{k,\sigma} \varepsilon_{k} c_{k\alpha}^{\dagger} c_{k\alpha} - 0^{+} S_{z} \\ &\text{infinitesimal magnetic field} \\ \Rightarrow \text{ Product initial state: } |\Psi_{i}\rangle &= |\uparrow\rangle \otimes |FS\rangle \\ H_{f} &= \sum_{k,\sigma} \varepsilon_{k} c_{k\alpha}^{\dagger} c_{k\alpha} - 0^{+} S_{z} + J \vec{S} \cdot \sum c_{k'\alpha} \vec{\sigma}_{\alpha\beta} c_{k\beta} \\ \text{ferromagnetic coupling (J<0): Coupling constant flows to zero} \\ &\Rightarrow \text{ Expansion becomes better} \\ &\text{(asymptotically exact) for long times} \\ \text{Nonequilibrium spin} \\ \text{expectation value} \\ (\text{large times}): & \langle S_{z}(t) \rangle = \frac{1}{2} \left(\frac{1}{\ln(t) - \frac{1}{\rho J}} + 1 + \rho J + O(J^{2}) \right). \\ \text{Equilibrium: } & \langle S_{z} \rangle_{eq} = \frac{1}{2} \left(1 + \frac{\rho J}{2} + O(J^{2}) \right) \end{split}$$





Crossover from adiabatic to instantaneous quenching: (C. Tomaras, S. K., Eur. Phys. Lett. 93, 47011 (2011))

Coupling J switched on on timescale $\boldsymbol{\tau}$

Measure of non-adiabacity:

$$\mu \stackrel{def}{=} \frac{\lim_{t \to \infty} \langle O(t) \rangle_{neq} - \langle O \rangle_0}{\langle O \rangle_{eq} - \langle O \rangle_0}$$

⇒ Crossover timescale nonperturbative (exponentially large) due to RG flow



5. Applications (bulk models): Quantum quench in a Fermi liquid



What happens for sudden switching (global quantum quench)?

Translation-invariant closed system + nonzero excitation energy density

 \Rightarrow Thermalization?

M. Moeckel and S. K., Phys. Rev. Lett. 100 (2008), Ann. Phys. 324 (2009)

Hubbard model in d>1 dimensions

$$H = \sum_{k,\alpha} \varepsilon_k c_{k\alpha}^{\dagger} c_{k\alpha} + U \Theta(t) \sum_i \left(n_{i\uparrow} - \frac{1}{2} \right) \left(n_{i\downarrow} - \frac{1}{2} \right)$$

Forward transformation:

$$c_{k\uparrow}^{\dagger}(B = \infty) = h_k c_{k\uparrow}^{\dagger} + \sum_{k'_1, k'_2, k_1} M_{k'_1 k'_2 k_1}^k : c_{k'_1\uparrow}^{\dagger} c_{k'_2\downarrow}^{\dagger} c_{k_1\downarrow} : + \text{ higher order terms}$$

$$\bullet h_k \text{ only nonzero at Fermi surface for zero temperature}$$

• Quasiparticle residue (equilibrium)

$$Z = h_{k_F}^2$$

Real time evolution

Analytical evaluation (with numerical integration) up to order U²





Numerical confirmation:

Non-equilibrium DMFT with real time QMC M. Eckstein et al., Phys. Rev. Lett. 103 (2009)

Sudden switching looks like T=0 Fermi liquid with "wrong" quasiparticle residue on time scale t $\propto D^{-1}$

- → Novel metastable prethermalized state
- → Thermalization via QBE on time scale t \propto U⁻⁴



5. Applications (bulk models): 1d fermions with dimerization

F. Essler, S. K., S. Manmana and N. Robinson, arXiv:1311.4557, to appear in PRB

$$H(\delta, U) = -J \sum_{l} \left(1 + (-1)^{l} \delta \right) \left(c_{l}^{\dagger} c_{l+1} + \text{h.c.} \right) + U \sum_{l} c_{l}^{\dagger} c_{l} c_{l+1}^{\dagger} c_{l+1}$$

 $H(\delta,U=0)$: Peierls insulator, exactly solvable via Bogoliubov transformation, exactly solved by flow equations

H($\delta \neq 0, U \neq 0$): Non-integrable model

Goal: Thermalization for global quenches $H(\delta_i, U_i=0) \rightarrow H(\delta_f, U_f \neq 0)$? (Here: tuneable integrability breaking $U_f \neq 0$ for fixed quench amplitude $\delta_i \rightarrow \delta_f$)

Flow equation calculation: - Up to terms $O(U^2)$

- Analytical calculation (with numerical integration)
- 2- and 4-point functions
- comparison with t-DMRG





FIG. 10. Comparison of the CUT and t-DMRG results for $\mathcal{G}(L/2, L/2 + 1) = \langle c_{L/2} c_{L/2+1}^{\dagger} \rangle$ for the quench $\delta_i = 0.75 \rightarrow \delta = 0.5$ and $U_i = 0 \rightarrow U = 0.15$ on a L = 50 chain. The

FIG. 12. Comparison of the CUT and t-DMRG results for $\mathcal{G}(L/2, L/2 + 1) = \langle c_{L/2} c_{L/2+1}^{\dagger} \rangle$ for the quench $\delta_i = 0.75 \rightarrow \delta = 0.5$ and $U_i = 0 \rightarrow U = 0.5$ on a L = 50 chain.





FIG. 13. Rescaled difference between the t-DMRG and CUT data for $\mathcal{G}(25, 26)$ and different values of U.

FIG. 18. Nearest neighbour density-density correlation function $\langle n(\frac{L}{2})n(\frac{L}{2}+1)\rangle$ for a quench from $\delta_i = 0.8 \rightarrow \delta_f = 0.4$ and $U = 0 \rightarrow 0.4$ computed by t-DMRG for system size

6. Outlook

Unitary perturbation theory based on flow equations:

- Perturbative method in the sense of weak coupling RG
- No secular terms in time evolution
- Gives exponential/power-law/etc. decays in lowest order calculations
- How to incorporate Boltzmann equation dynamics?
- Convergence properties of observable transformation?